#### **CHAPTER 12**

#### RADIATION

### §12.01 General considerations

There are several alternative ways of approach to the thermodynamics of radiation. We shall choose the one according to which the radiation is regarded as a collection of photons. Each photon is characterized by a frequency, a direction of propagation, and a plane of polarization. In empty space all photons have equal speeds c. Each photon has an energy  $U_i$  related to its frequency  $v_i$  by Planck's relation

$$U_i = hv_i 12.01.1$$

and a momentum of magnitude  $hv_i/c$ . It is convenient to group together all the species of photons having equal frequencies, and so equal energies, but different directions of propagation and planes of polarization. We denote by  $g_i$  the number of distinguishable kinds of photons having frequencies  $v_i$  and energies  $U_i$ . More precisely  $g_i dv_i$  denotes the number of distinguishable kinds of photons having frequencies between  $v_i$  and  $v_i + dv_i$  and energies between  $U_i$  and  $U_i + dU_i$ . By purely geometrical considerations it can be shown\* that in an enclosure of volume V

$$g_i dv_i = 2 \times 4\pi V c^{-3} v_i^2 dv_i$$
 12.01.2

the factor 2 being due to the two independent planes of polarization.

# §12.02 Energy and entropy in terms of $g_i$ 's

We denote by  $N_i$  the number of photons having energy  $U_i$  and frequency  $v_i$  interrelated by (12.01.1). Then the total energy U is given by

$$U = \sum_{i} N_{i} U_{i}.$$
 12.02.1

<sup>\*</sup> Brillouin, Die Quantenstatistik, Springer 1931 ch. 2; Fowler and Guggenheim, Statistical Thermodynamics, Cambridge University Press 1939 §§ 401-403.

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From the fact that photons obey Bose-Einstein statistics it can be shown\* that the entropy S of the system is given by

$$S/k = \sum_{i} \ln\{(g_i + N_i)!/g_i! N_i!\}.$$
 12.02.2

Differentiating (1) and (2) at constant  $g_i$ , that is to say constant V, we have

$$dU = \sum_{i} U_{i} dN_{i}$$
 12.02.3

$$dS/k = \sum_{i} \ln\{(g_i + N_i)/N_i\} dN_i.$$
 12.02.4

The condition for equilibrium is according to (1.35.1) that S should be a maximum for given U, V. Hence for the most general possible variation, the expressions (3) and (4) must vanish simultaneously. It follows that

$$U_i/\ln\{(g_i+N_i)/N_i\} = U_k/\inf\{(g_k+N_k)/N_k\}$$
 (all  $i, k$ ) 12.02.5

and consequently using (3) and (4)

$$U_i/\ln\{(g_i + N_i)/N_i\} = \sum_i U_i dN_i / \sum_i \ln\{(g_i + N_i)/N_i\} dN_i$$
$$= k dU/dS = kT$$
 12.02.6

since at constant volume

$$dU = T dS \qquad (V \text{ const.}). \qquad 12.02.7$$

From (6) we have

$$N_i/(g_i + N_i) = \exp(-U_i/kT)$$
 12.02.8

and so

$$N_i = g_i / \{ \exp(U_i / kT) - 1 \}.$$
 12.02.9

Substituting (9) into (1), we obtain

$$U = \sum_{i} g_{i} U_{i} \{ \exp(U_{i}/kT) - 1 \}.$$
 12.02.10

For the entropy we obtain from (2), using Stirling's approximation for the factorials, and by use of (8)

$$S = \sum_{i} N_{i} \ln\{(g_{i} + N_{i})/N_{i}\} + \sum_{i} g_{i} \ln\{(g_{i} + N_{i})/g_{i}\}$$

$$= \sum_{i} N_{i} U_{i}/kT - \sum_{i} g_{i} \ln\{1 - \exp(-U_{i}/kT)\}.$$
12.02.11

For the Helmholtz function F we deduce from (1) and (11)

$$AF = kT \sum_{i} g_{i} \ln\{1 - \exp(-U_{i}/kT)\}.$$
 12.02.12

<sup>\*</sup> Brillouin, Die Quantenstatistik, Springer 1931 ch. 6.

# §12.03 Thermodynamic functions

In the previous section we obtained formulae for the energy, the entropy, and the Helmholtz function in terms of the  $U_i$ 's and  $g_i$ 's without making any use of (12.01.1) or (12.01.2). If we now substitute the values of  $U_i$  and  $g_i$ , given by these formulae, into the relations of the previous section we obtain

$$\mathcal{F} = 8\pi V c^{-3} kT \int_0^\infty \ln\{1 - \exp(-h\nu/kT)\} v^2 d\nu$$
 12.03.1

$$U = 8\pi V c^{-3} \int_0^\infty h v^3 \{ \exp(hv/kT) - 1 \}^{-1} dv.$$
 12.03.2

We can write (2) in the form

$$U = \int_0^\infty U_{\nu} d\nu \qquad 12.03.3$$

$$U_{\nu} = 8\pi V c^{-3} h v^{3} \{ \exp(hv/kT) - 1 \}^{-1}$$
 12.03.4

which is Planck's formula from which quantum theory originated.

# §12.04 Evaluation of integrals

We can rewrite (12.03.1) as

$$\mathcal{F} = -8\pi V k^4 T^4 h^{-3} c^{-3} I$$
 12.04.1

where I is the integral defined by

$$I = -\int_{0}^{\infty} \xi^{2} \ln\{1 - \exp(-\xi)\} d\xi.$$
 12.04.2

Using the power series for the logarithm and then integrating term by term, we obtain

$$I = \int_0^\infty \sum_{n=1}^\infty n^{-1} \xi^2 \exp(-n\xi) d\xi = \sum_{n=1}^\infty n^{-4} \int_0^\infty \eta^2 \exp(-\eta) d\eta$$
$$= 2 \sum_{n=1}^\infty n^{-4} = \pi^4 / 45.$$
 12.04.3

Substituting (3) into (1) we obtain finally

$$\mathcal{A} = -(8\pi^5 k^4 / 45c^3 h^3) T^4 V.$$
 12.04.4

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### §12.05 Stefan-Boltzmann law

We could obtain formulae analogous to (12.04.4) for U and S by evaluation of the relevant integrals, but it is more convenient to obtain these formulae by differentiation of (12.04.4).

We first abbreviate (12.04.4) to

$$AF = -\frac{1}{3}aT^4V$$
 12.05.1

where a is a universal constant defined by

$$a = 8\pi^{5}k^{4}/15c^{3}h^{3} = 7.5646 \times 10^{-16} \text{ J m}^{-3} \text{ K}^{-4}.$$
 12.05.2

From (1) we deduce immediately

$$S = \frac{4}{3}aT^3V$$
 12.05.3

$$U = aT^4V$$
 12.05.4

$$P = \frac{1}{3}aT^4 = \frac{1}{3}U/V$$
 12.05.5

$$G = U - TS + PV = 0.$$
 12.05.6

Formula (5) can be derived from classical electromagnetic theory. Formula (4) was discovered by Stefan and derived theoretically by Boltzmann. It is called the *Stefan-Boltzmann law*.

From (4) we see that  $aT^4$  is the equilibrium value of the radiation per unit volume in an enclosure. If a small hole is made in such an enclosure then it can be shown by geometrical considerations that the radiation emitted through the hole per unit area and per unit time is  $\sigma T^4$ , where  $\sigma$  is given by

$$\sigma = \frac{1}{4}ac = 5.670 \times 10^{-8} \text{ J m}^{-2} \text{ s}^{-1} \text{ K}^{-4}$$
 12.05.7

in which c denotes the speed of light. This constant  $\sigma$  is called the Stefan-Boltzmann constant.

# §12.06 Adiabatic changes

Suppose that radiation is confined by perfectly reflecting walls and that the volume of the container is altered by moving a piston. If the radiation remains in thermal equilibrium its temperature will change. For such a reversible adiabatic change, we have

$$S = \text{const.}$$
 12.06.1

From (12.05.3) and (1) it follows that

$$VT^3 = \text{const.}$$
 (adiabatic). 12.06.2

From (12.05.4) and (12.05.5) we have

$$P/T^4 = \text{const.}$$
 12.06.3

so that

$$PV/T = \text{const.}$$
 (adiabatic) 12.06.4

and

$$PV^{\frac{4}{3}}$$
 = const. (adiabatic). 12.06.5

From (2), (3), (4), (5) it appears that the relations for a reversible adiabatic change in radiation are formally similar to those for a perfect gas such that the ratio  $C_P/C_V$  has the constant value  $\frac{4}{3}$ . This apparent resemblance is however accidental, for the ratio  $C_P/C_V$  of radiation is not  $\frac{4}{3}$ . In fact for radiation

$$C_V = (\partial U/\partial T)_V = T(\partial S/\partial T)_V = 4aT^3V$$
 12.06.6

while

$$C_{\mathbf{p}} = T(\partial S/\partial T)_{\mathbf{p}} \to \infty$$
 12.06.7

since no increase in S, however great, can increase T without increasing P.