

Solar Radiation:-

- * The power of the sun @intercepted by earth is 1.8×10^{11} MW. and if we harness this energy with the 0.01% efficiency it is sufficient for world power consumption.
- * The solar radiation flux having a maximum value of 1 kW/m^2 and over a day it is maximum of 7 kWh/m^2 .

Solar energy Quantification:-

$$D_s = 1.39 \times 10^6 \text{ km} \quad \text{Diameter of Sun}$$

$$D_e = 1.27 \times 10^4 \text{ km} \quad \text{Diameter of earth}$$

$$L_m = 1.5 \times 10^8 \text{ km} \quad \text{mean distance between sun & earth}$$

Frequency \propto temp

$$\nu \propto T$$

$$\lambda \propto \frac{1}{\nu}$$

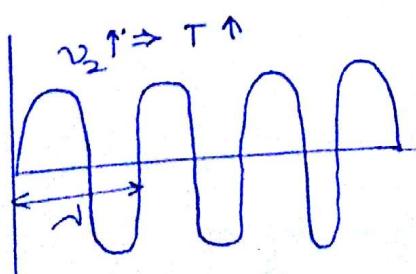
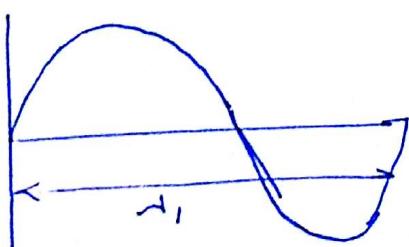
$$\lambda \cdot \nu = C$$

$$\Rightarrow \nu \propto T \propto \frac{1}{\lambda}$$

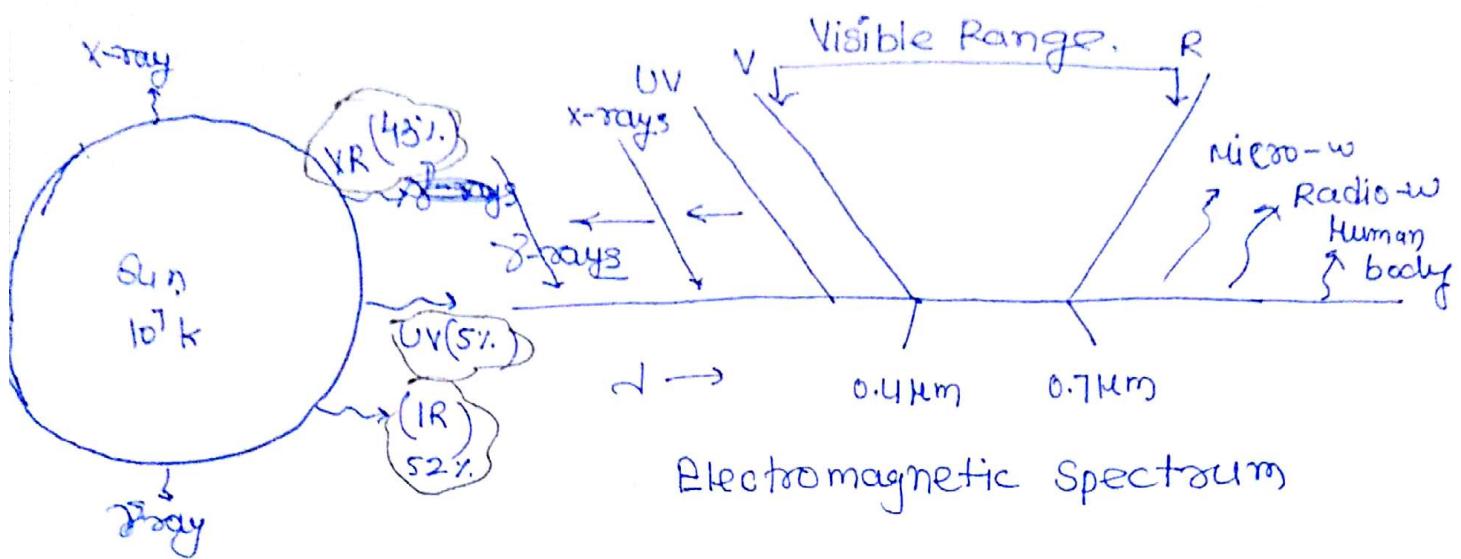
$$\Rightarrow \lambda \cdot \nu = C (3 \times 10^8 \text{ m/s})$$

$$\lambda \cdot T = \text{const.} (2898 \text{ nm}\cdot\text{K})$$

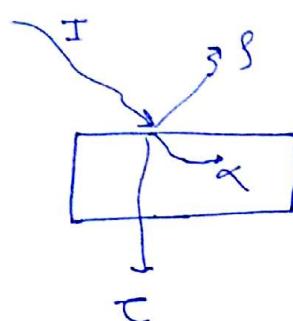
$$\frac{\nu}{T} = \frac{C}{2898}$$



More penetration



Electromagnetic Spectrum



$$\alpha + \gamma + \epsilon = 1$$

short-wave

$\alpha \rightarrow E$

$\gamma, \text{UV+IR+VR}$

E

This diagram shows the law of radiation. It states that the sum of the fractions of energy emitted by a body as short-wave radiation (α), long-wave radiation (γ), and thermal radiation (ϵ) equals 1. The term α is associated with the Earth (E).

Fig:- Basic surface property

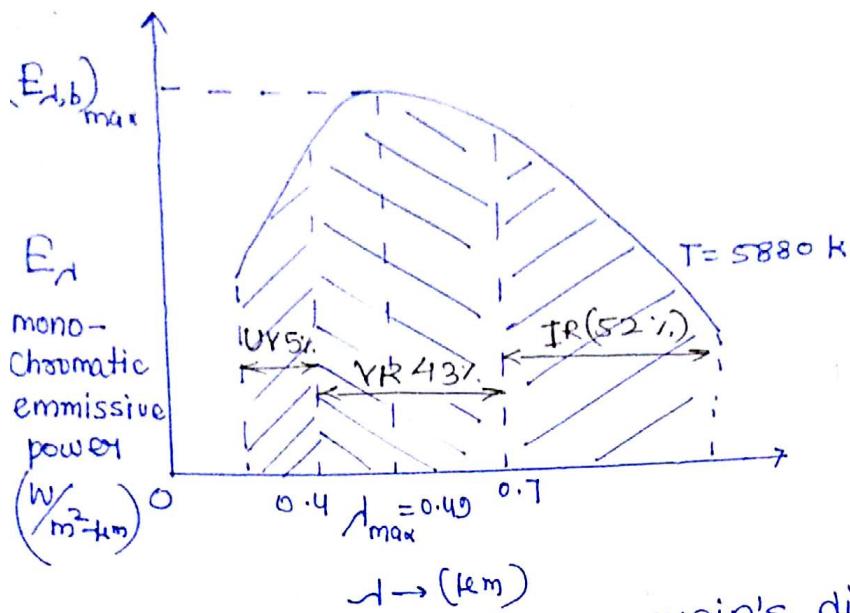
UV - Ultra violet. (5%)

VR - Visible range (43%)

IR - Infrared (52%)

Electromagnetic Spectrum of Solar Radiation: —

- * Solar radiation covers the entire range of visible region and some part of UV & IR waves. Th.
- * The maximum emissive power reaching the earth surface (extra-terrestrial solar flux) is in the Infrared region but the maximum monochromatic spectral emissive power is in the visible region ($\lambda_{\max} = 0.49 \mu\text{m}$)



Plank's law

$$E_{d,b} \propto \frac{1}{e^{\frac{C_2}{\lambda T}}}$$

Wein's displacement law

$$E_{d,b} \propto \frac{1}{e^{\frac{C_2}{\lambda T}}}$$

$$\lambda_{\max} \cdot T = 2898 \text{ km} \cdot \text{K}$$

(For black & Gray body
not valid for real surface)

Stefan-Boltzmann law

$$E_b = \int_0^{\infty} E_{d,b} \cdot d\lambda = \sigma T^4$$

~~Kelvin~~ ~~law~~ Kirchhoff's law
At thermal eq^m
 $\alpha = \epsilon$

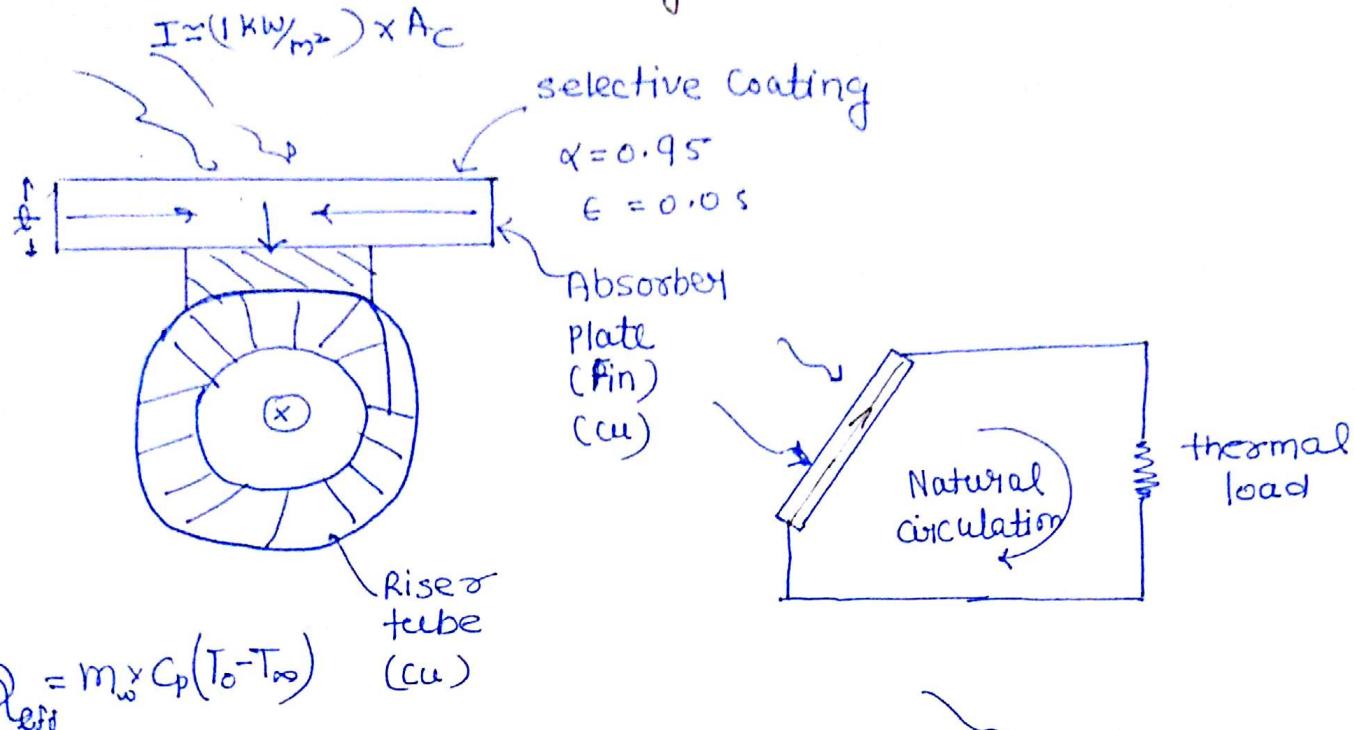
Question Determine the range of wavelength for which earth surface is emitting max. spectral emissive power if the earth temp range from 250 to 300 K

$$\text{Sol}^9 \quad \lambda_{m_1} \cdot 250 = 2898 \Rightarrow \lambda_{m_1} = 11.592 \text{ km}$$

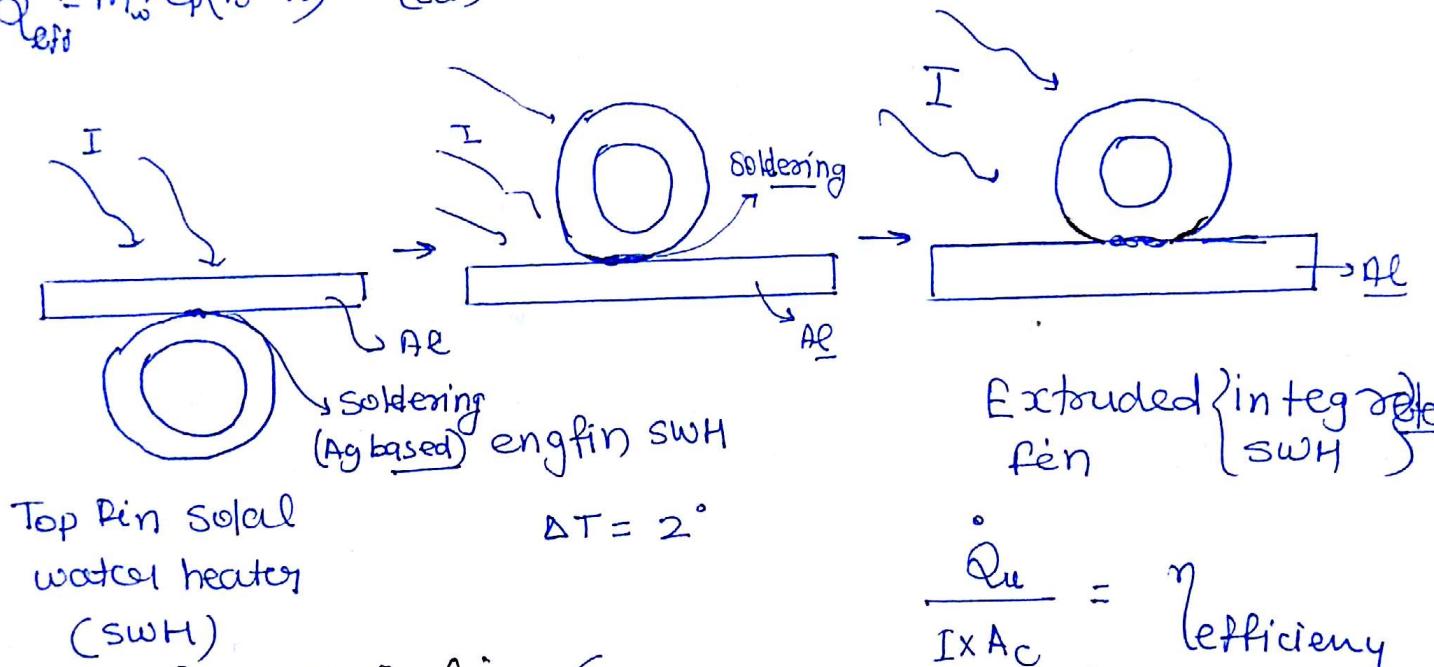
$$\lambda_{m_2} \cdot 300 = 2898 \Rightarrow \lambda_{m_2} = 9.66 \text{ km}$$

So the range of wave length in which earth emit max. spec. emissive power is $\lambda_m \in (11.59, 9.66) \text{ km}$

Solar water Heater Analysis:-



$$Q_{\text{eff}} = m_w \times C_p (T_o - T_{\infty})$$

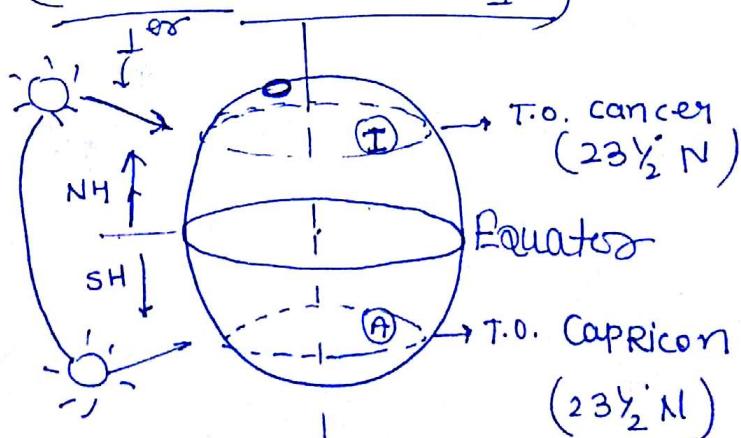
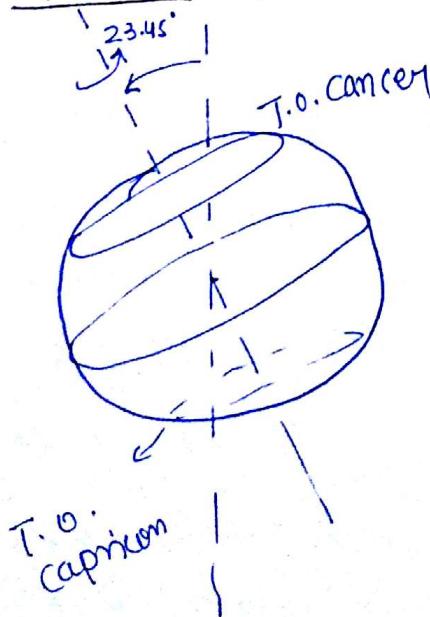


Top Fin Solar water heater (SWH)

$$\Delta T = 2^\circ$$

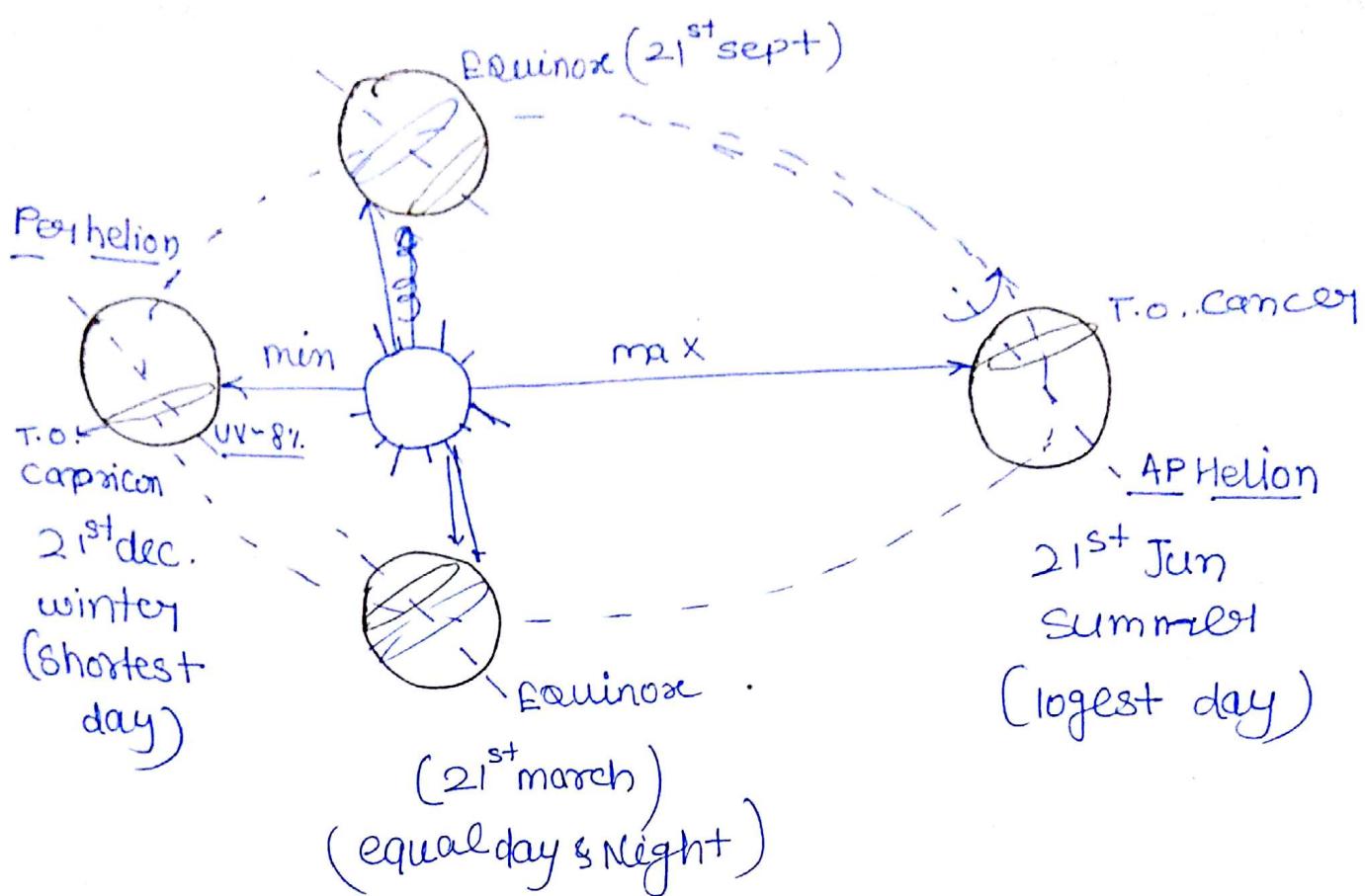
$$\frac{Q_u}{I \times A_C} = \eta_{\text{efficiency}}$$

Solar flux Variation ($I_{sc} = \text{Solar Const. } \perp^2$)



NH - Northern Hemisphere

SH - Southern Hemisphere



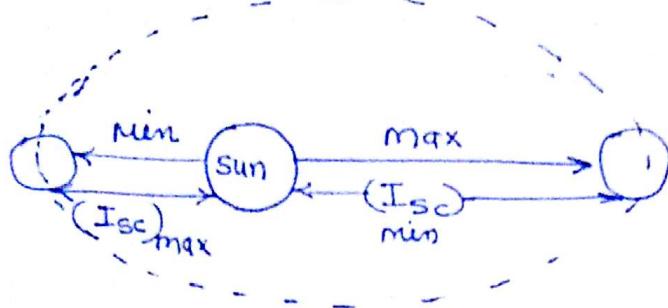
The orbit on which earth revolves around the sun is elliptical hence the distance between the earth and sun varies w.r.t. day.

The maxⁿ solar flux strike the earth surface is on 21st Dec. in Southern hemisphere whereas the min on 21st Jun in Northern hemisphere (NH) in 1^{er} direction.

Solar Constant:- The energy flux receive from the sun to the outside surface of earth.

$$I_{sc} = 1367 \text{ W/m}^2 \rightarrow \text{Solar constant.}$$

* The orbit of the earth on which earth revolves around the sun is elliptical hence the mean distant b/w sun and earth varies and so thus the solar constant also changes.



$$I_{sc} = 1367 \text{ W/m}^2$$

total variation of I_{sc} = $\pm 3.3\%$.

for n^{th} day

$$I'_{sc} = I_{sc} \left\{ 1 + 0.033 \cos \left(360 \times \frac{n}{365} \right) \right\}$$

$$\cos \theta \begin{cases} \text{Max} = +1 \\ \text{Min} = -1 \end{cases} \quad n \rightarrow n^{th} \text{ day}$$

$$n=1 \Rightarrow 1 \text{ Jan.}$$

$$\text{For Jun 21, } n = 31 + 28 + 31 + 30 + 31 + 21 = 172$$

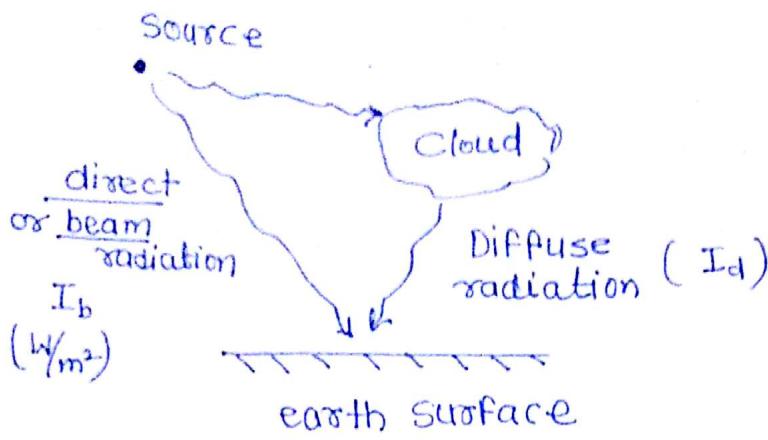
$$I'_{sc} = 1320.6 \text{ W/m}^2$$

$$\text{For Dec 21, } n = 355$$

$$I'_{sc} = 1491.4 \text{ W/m}^2$$

Effect of Environment on Solar radiation:-

Solar radiation received on the earth surface are undergone Absorption, scattering or reflection from the ozone layer and water vapour.



Total Radiation

$$I_g = \text{Global radiation} = I_b + I_d$$

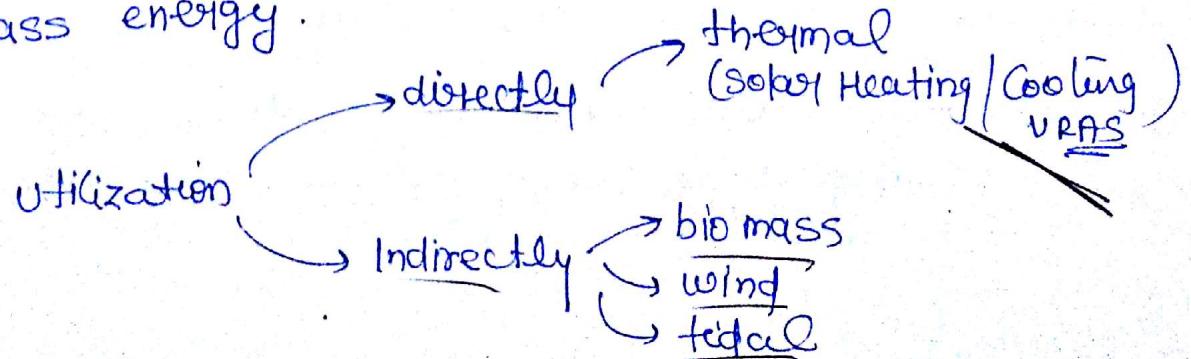
Direct/Beam Radiation:- Solar radiation receive at the surface of earth in line with the sun is known as direct radiation.

Diffuse Radiation:- Solar radiation receive on the earth after getting scatter because of the atmosphere of earth.

Utilization Method for solar Energy:-

Solar energy can be utilise directly or indirectly in which directly heat of water or air or conversion of solar into electricity.

Indirect method of solar energy derives the ecology system. Some of the examples are. wind energy and Bio-mass energy.



Directly :-

(A) Thermal - solar heating or cooling

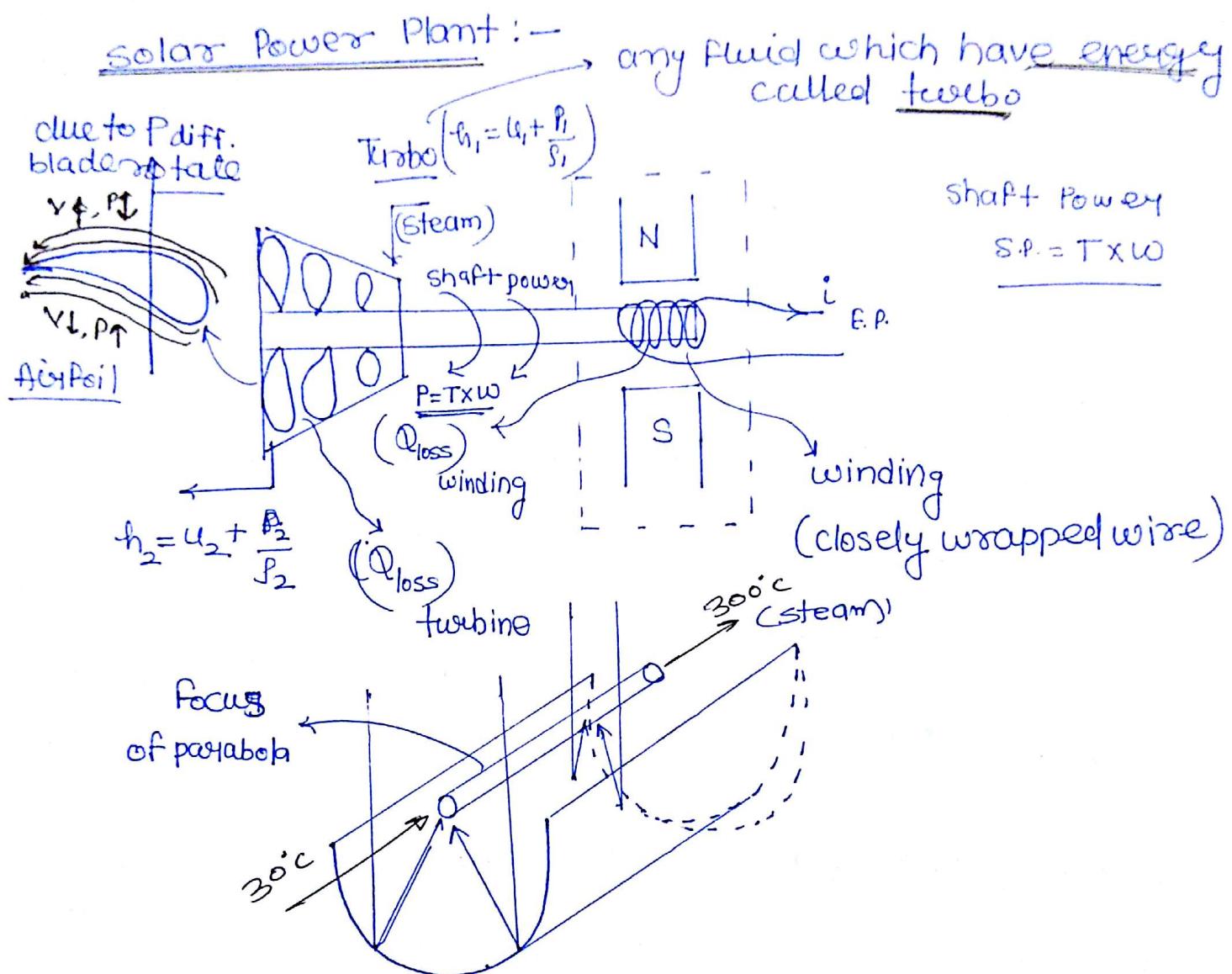


Fig: Parabolic Collector

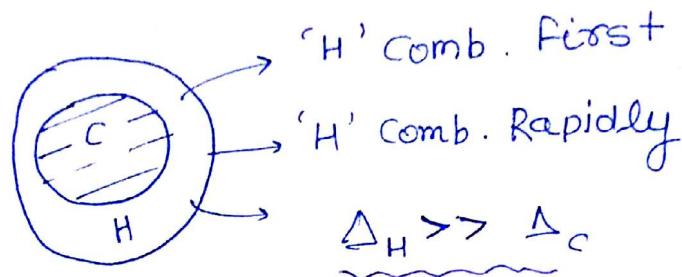
- * A high energy contentent fluid (turbo) is entered into the turbine and losses its energy and that energy is given the turbine shaft power
- * In order to create high energy fluid like steam Concentrating type solar collectors are used.

* (B) Photo-Voltaic

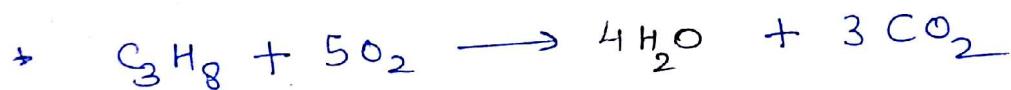


Indirectly

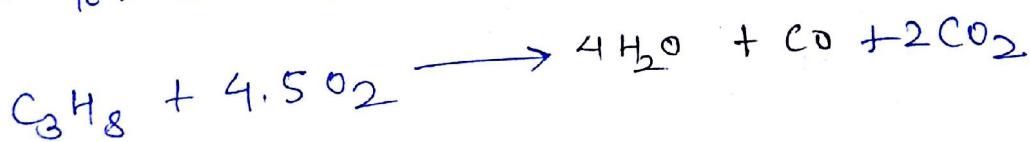
Bio mass :-



stoichiometry reaction / perfect combustion

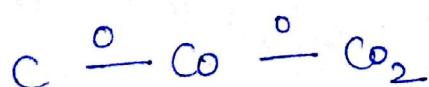


let 10% less O_2



$$\% \text{ CO} = \frac{1}{4+1+2} = 14.28\%$$

Gate-2016

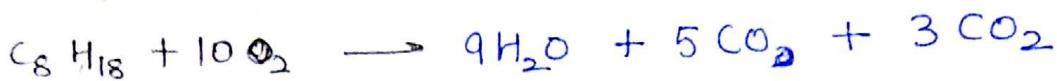


* First rapidly 'H' comb. happens and H_2O becomes then rest O_2 make CO and then CO_2

Complete
Combustion



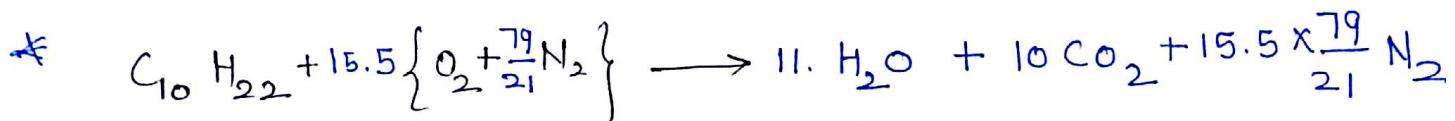
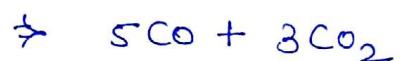
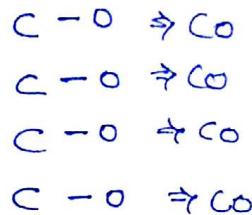
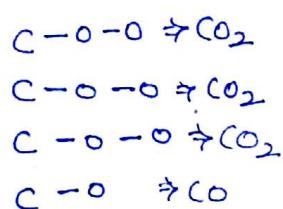
* Stab. let $10O_2$



$$\% CO = \frac{5}{9+5+3} = 29.4\%$$

पहले H_2O बढ़ाए $\xrightarrow{11 \rightarrow 0}$ left

फिर CO \Rightarrow
एक एक O से
और पिछे CO_2



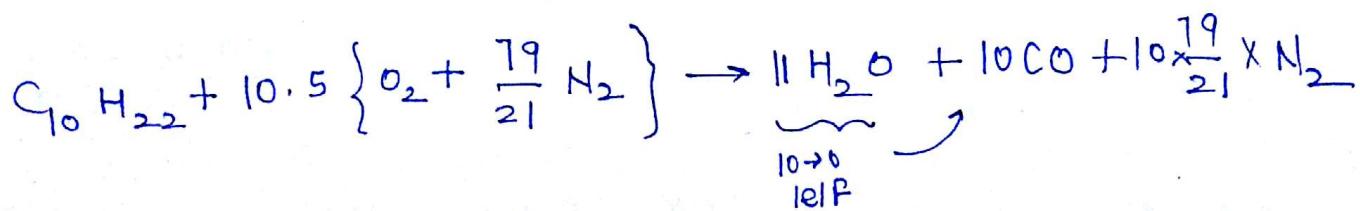
$$\begin{aligned} 27+20 \\ = 31 \\ \text{Required} \end{aligned}$$

$$\underline{\text{So } 15.2O_2}$$

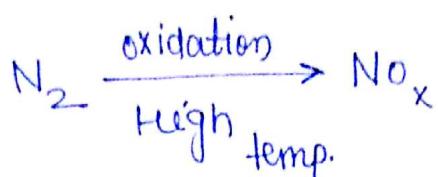
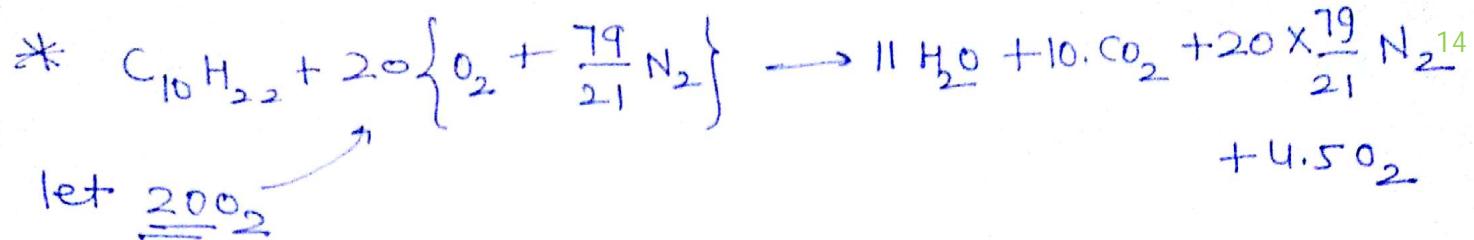
Air

$$\left. \begin{array}{l} 21\% \rightarrow O_2 \\ 79\% \rightarrow N_2 \\ \text{for } 1 O_2 \\ \frac{79}{21} N_2 \end{array} \right\}$$

let $10.5O_2$ then



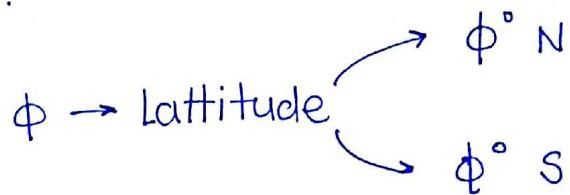
$$\text{So. } \% CO = \frac{10}{10+11+10 \times \frac{79}{21}} = 17.05\%$$



Latitude and longitude :- Definition of Angle

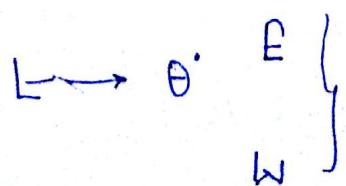
Latitude Angle (ϕ) :- The latitude of a location P is the angle made by the radial line joining the location to the center of earth with the projection of line on equatorial plane.

Sign Convention :-

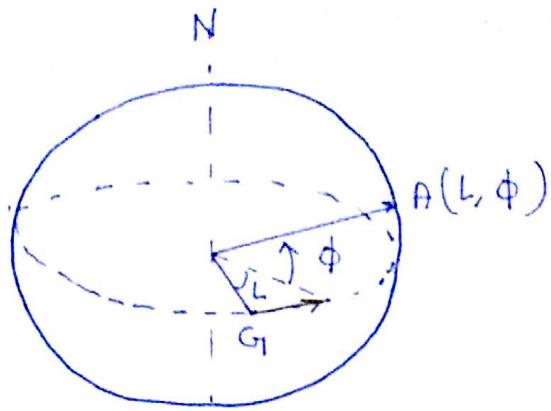
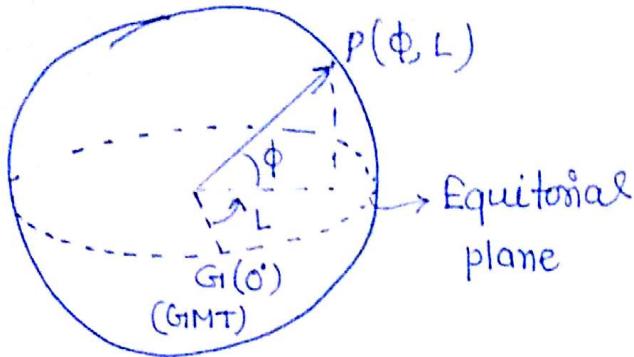


Longitude (L) :- longitude is measured from the reference point Greenwich, England and it is the angle between the projection of line of any point P on equatorial plane the projection of line having latitude of Greenwich.

sign convention

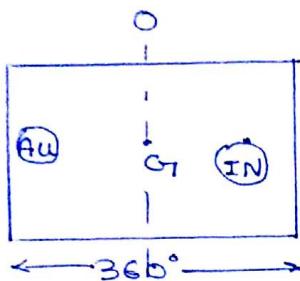


From Greenwich england



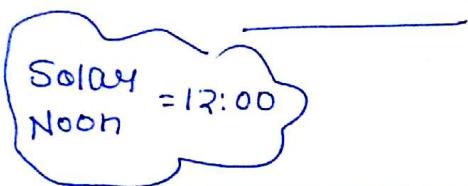
G₁ - Greenwich

G_{MT} - Global Meridian time
(Greenwich)

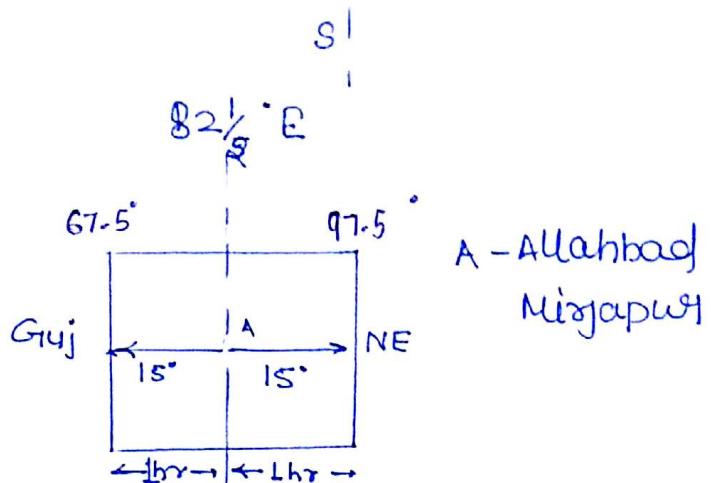


$$\text{one day} = 360^\circ$$

$$1 \text{ hr} = \frac{360}{24} = 15^\circ/\text{hr}$$



$$\boxed{\text{Solar Time} = \text{Watch time} - 4(L_{st} - L_{lo})}$$



Diff b/w Allahabad & Green which

$$= \frac{82\frac{1}{2}}{15} = 5.5 \text{ hrs}$$

(depend only on longitude)

$$\boxed{ST = WT - 4(L_{st} - L_{lo})}$$

L_{st} = longitude of standard Meridian (Country)

L_{lo} = local longitude

$$(ST)_{NE} = 12 - 4(-15) = 1300 \text{ hrs}$$

$$(ST)_{Guj} = 12 - 4(15) = 1100 \text{ hrs}$$