GRAVITATION

• Newton's Law of Gravitation: Every particle in the universe attracts every other particle with a force. The force of attraction between two masses is directly proportional to the product of their masses and inversely proportional to the square of the distance between them.

$$F = G \frac{m_1 m_2}{d^2}$$

where m_1 and m_2 are masses of two particles dis the distance of separation between their centres and G is universal gravitational constant Value

of $G = 6.67 \times 10^{-11} Nm^2 kg^{-2}$ (or) 6.67×10^{-8} dyne $cm^2 g^{-2}$

- Gravitational foce acts along the line joining the centres of the bodies.
- Gravitational force between two particles is independent of the properties of intervening medium and the presence of other bodies.
- The value of G is same for a) all pairs of bodies b) all types of media
- G is a scalar quantity with dimensional formula $M^{-1}L^3T^{-2}$
- Kepler's laws can be deduced from Newton's law of gravitation.
- If two identical spheres each of radius 'r' made up of same material are kept in contact with each other, the gravitational force acting between them

$$F = \frac{Gm_1m_2}{(2r)^2} = \frac{G\left[\frac{4}{3}\pi r^3\rho\right]^2}{(2r)^2}$$

 $\Rightarrow F \alpha r^4$

- If two identical spheres each of radius 'r' made up of same material are separated by a constant distance then $\Rightarrow F \propto r^6$
- If two bodies of masses m₁ and m₂ are separated by a distance 'd' then the distance to the null point 'P' from m₁ is

$$x = \frac{d}{\sqrt{\frac{m_2}{m_1} + 1}}$$

(a) Gravitational force is always attractiveb) It is a conservative force. Workdone by it is path independent. The workdone in moving a particle around a closed path under the action of gravitational force is zero.

c) If two particles of equal mass 'm' are placed at the two vertices of an equilateral triangle of side 'a', then the resultant gravitational force on unit mass placed at the third vertex is given by

$$F = \sqrt{F_1^2 + F_2^2 + 2F_1F_2\cos\theta}$$
$$= \sqrt{3}F \quad [Q \ F_1 = F_2]$$
$$\sqrt{3}\left[\frac{Gm}{a^2}\right]$$

a) Range of gravitational force> Range of electromagnetic force> Range of nuclear force
b) Strength of nuclear force> strength of electromagnetic force> strength of gravitational force.
Relation between g & G

a) The acceleration due to gravity 'g' on the surface of a planet of mass M and radius R is

$$g = \frac{GM}{R^2}$$

•

where G = universal gravitational constant

b)
$$g = \frac{4}{3}\pi R\rho G$$

where ρ is density of the planet

Orbital velocity

The velocity of a satellite revolving around earth of mass M and radius R in a circular orbit of radius 'r' at a height 'h' from the surface of earth is called orbital velocity.

$$V_o = \sqrt{\frac{GM}{r}} = \sqrt{\frac{GM}{(R+h)}} = \sqrt{\frac{gR^2}{(R+h)}}$$

when
$$h \ll R$$
 then, $V_o = \sqrt{\frac{GM}{R}} = \sqrt{gR}$

For two satellites revolving around the earth in different cicular orbits of radii r_1 and r_2 at vertical heights h_1 and h_2

$$\frac{V_1}{V_2} = \sqrt{\frac{r_2}{r_1}} = \sqrt{\frac{R+h_2}{R+h_1}}$$

- Orbital velocity $V_0 = 7.92 \text{ kms}^{-1} \approx 8 \text{ kms}^{-1}$ (close to the surface of earth)
- It is independent of mass of the satellite and angle of its projection. It is always along the tangent to the orbit.
- Close to the surface of planet, $V_e = \sqrt{2} V_0$
- If the speed of the orbiting body is made $\sqrt{2}$ times its initial velocity or velocity is increased by 41% or its KE is doubled, then the body will escape.
- If the speed of the orbiting body 'V' is such that
 V_o < V < V_e then its orbit changes from circle to ellipse

• Time period of revolution
$$T = \frac{2\pi}{\omega} = \frac{2\pi r}{V_0}$$

$$\Rightarrow T = \frac{2\pi r}{R} \sqrt{\frac{r}{g}}$$

• Time peirod is independent of mass of the orbiting body and depends on mass of the planet and radius of the orbit.

•
$$T^2 = 4\pi^2 \frac{r^3}{GM} \Rightarrow \frac{T^2}{r^3} =$$
 Constant, which is

Kepler's third law of planetary motion.

• If a satellite is revolving close to the surface of the earth, then the radius of the orbit is taken as the radius of earth.

$$\therefore T = 2\pi \sqrt{\frac{R}{g}} = 84.6_{\text{min}}$$

- If a satellite revolves round the earth from west to East in the equatorial plane with a period equal to the period of rotation of earth, then it appears to be stationary. This orbit is known as 'Parking Orbit' and the satellite is known as 'Geostationary Satellite'
- The velocity of geo-stationary satellite relative to earth is zero.
- The velocity of geo-stattionary satellite is 3.1 kms^{-1}
- Radius of parking orbit is 42,400 km.
- The height of the parking orbit above the surface

of the earth is 36,000 km

- Escape velocity
- The velocity required for a body to escape into space from earth's gravity is called escape velocity.

$$v_e = \sqrt{\frac{2GM}{R}} = \sqrt{2gR} = \sqrt{2G\left[\frac{4\pi R^3 \rho}{3R}\right]}$$
$$= 2R\sqrt{\frac{2}{3}G\pi\rho} \quad \text{where } \rho \text{ is density of earth}$$

• From earth's surface, $V_e = 11.2 \text{ Kms}^{-1}$

From the surface of the moon $V_e = 2.38$ Kms⁻¹ Escape velocity is independent of mass of the projected body and angle of its projection.

- Escape velocity on a planet of mass M_1 and radius
 - R_1 is given by

$$V_e^1 = V_e \sqrt{\left[\frac{M_1}{M}\right] \left[\frac{R}{R_1}\right]}$$
$$\Rightarrow V_e^1 = V_e \left[\frac{R_1}{R}\right] \left[\frac{\rho_1}{\rho}\right]$$

Where ρ_1 is density of the planet

- A planet will have atomosphere if the velocity of the molecules (C_{rms}) in its atmosphere is less than escape velocity
- There is no atmosphere on the surface of the moon since the escape velocity on the surface of the moon is less.
- Sufficient amount of hydrogen is present in the atmosphere of the sun since the escape velocity on the sun is very high.
- If a body is projected with a velocity less than the value of escape velocity $(V < V_e)$ it will reach a certain height and then either may move in an orbit around the planet or may fall back to the planet.
- If a body is projected with a velocity greater than the escape velocity $(V > V_e)$ then by conservation of energy

$$\frac{1}{2}mV^2 - \frac{GMm}{R} = \frac{1}{2}mV_1^2$$

 $\Rightarrow V_1^2 = V^2 - \frac{2GM}{R} = V^2 - V_e^2 \left[Q V_e^2 = \frac{2GM}{R} \right]$ $\Rightarrow V_1 = \sqrt{V^2 - V_e^2}$ i.e., the body will move in interplanetary or interstellar space with a velocity $\sqrt{V^2 - V_e^2}$ **Energy of statellite :** Gravitational potential is the work done in moving unit mass from the point to infinity against gravitational force. $V = \frac{-GM}{r}$ Gravitational potential Energy (GPE) of a body on the surface of the earth $U = \frac{-GMm}{R}$ M = mass of earthwhere m = mass of the bodyR = radius of earth Gravitational potential energy of a body at an altitude 'h' is $U = \frac{-GMm}{(R+h)}$ the work done or change in potential energy when a body of mass 'm' is displaced from a height h_1 to h_2 , with respect to the surface of earth is $W = \Delta PE = GMm \left| \frac{1}{(R+h_1)} - \frac{1}{(R+h_2)} \right|$ $= mgR^2 \left| \frac{1}{R+h} - \frac{1}{R+h_2} \right|$

• If a body of mass m' is displaced from the surface of earth to infinity then the change in its potential energy is

$$\Delta PE = mgR$$

where R is radius of earth

- Potential Energy at the centre of the earth is minimum but not zero.
- For a satellite orbiting close to the surface of earth

$$(i) KE = \frac{1}{2} \frac{GMm}{R}$$

$$ii)PE = -\frac{GMm}{R}$$

$$(iii) KE = \frac{PE}{2}$$
 (or) $PE = 2(KE)$

numerically.

$$(iv)$$
 Total Energy = $KE + PE = \frac{-GMm}{2R}$

- A surface satellite possesses both KE and PE
 - *i*) For a satellite in the orbit

(i) PE > KE

(*ii*) The energy to be supplied to it to escape into space is called binding energy (BE)

$$BE = \frac{GMm}{R}$$

- When the total energy of a satellite is zero, it escapes into space in a parabolic path.
- When a body escapes into space, its KE = PE
- The KE required to make a body escape into space is

$$KE = \frac{GMm}{R} = mgR$$

- When a satellite is lifted from a lower orbit to higher orbit then a) its PE increases and b) KE decreases
- The amount of workdone in lifting a body from the surface of the earth to a height 'h' is

$$W = \frac{GMm}{R\left[1 + \frac{R}{h}\right]}$$

• Geostationary satellites:

• Conditions for geo-stationary satellite:

• The plane of obit of the satellite should coincide with geo-equatorial plane

ii) The direction of revolution of satellite should be same as the direction of rotation of earth (ie., from West to East)

iii) The time period of revolution of the satellite should be 24 hr

Time period of revolution of geo-stationary satellite with respect to earth is infinity

• When a satellite is revolving around the earth	field from the centre of mass of the body, then
in a fixed orbit then the astronant inside the satel-	$E_{G} = \frac{GM}{r^{2}}$
lite experiences weightlessness or zero apparent	$E_{\rm G} = \frac{1}{r^2}$
weight because zero normal reaction is exerted by the satellite on the astronaut and both have	c) Units of gravitational field strength are Nkg^{-1}
equal acceleration towards the centre of the earth.	or ms^{-2} and dimensional formula is LT^{-2}
 The shape of the orbit of a satellite depends on 	d) It is a vector quantity, It is always directed ra- dially towards the centre of mass of the body pro-
its velocity $V_{.}$ Let V_{o} be the orbital velocity for	ducing the field.
a circular orbit	Note :- In the earth's gravitational field, the acceleration
<i>i</i>) If $V < V_o$ the satellite falls to the ground like a	due to gravity 'g' may be considered as the gravi- tational field strength.
projectile in a parabolic path	• Propagation of gravitational field :-
<i>ii</i>) If $V = V_o$ then the orbit is circular	a) According to Einstein's general theory of rela-
<i>iii</i>) If $V > V_{q}$ but $V < V_{e}$ then the orbit is ellipti-	tivity, whenever a body with mass is accelerated,
cal	the gravitational field around it undergoes rapid changes.
iv) If $V = V_e$ then the orbit is parabolic	b) Just as photon in electromagnetic field, a quan-
• If $V > V_e$ then the orbit is hyperbolic	tum of energy is associated with gravitational field called 'graviton'.
• Uses of Geo-sationary satellites:	c) gravitons, like photons would be mass less,
(i) to photograph different regions on earth's	electrically uncharged particles, travelling at the speed of light and would be emitted by highly ac-
surface	celerating and extremely massive objects such as
(ii) for weather forecast	d) Gravitational fields are propagated with the help
(<i>iii</i>) for telecommunication, television transmission	of gravitons with speed of light.
	Black Hole :
(iv) to locate the places of mineral deposits on	a) During the dying stage of a massive star, supernova is formed.
the earth.	b) If the core of supernova exceeds about two
(v) for spying purpose	solar masses, it continues to contract.
(vi) to study different experiments in weight-	c) The gravitational field of the collapsing star is predicted to be so powerful that neither matter
less conditions.	nor light can escape it.
Note : A minimum of 3geo-sationary satellites at an	d) The "star" then collapses to a black hole
angular separation of 120° are required for live	e) It is a point of zero volume and infinite mass.f) It is hidden by an event horizon at a distance
telecast of a programme all around the earth.	called Schwarzschild radius.
• Gravitational Field :-	g) Black holes remain hypothetical, but observa-
a) The concept of gravitational field is used to overcome the difficulties encountered in universal	tions suggest that such phenomena may possibly
law of gravitation.	exist in the star system Cygnus -X-1 and center of our Galaxy.
b) Einstein considered gravitational field as a dis-	 Schewarzschild Radius :
tortion of 'space' due to the presence of matter.	
• Gravitational field strength :- a) It is defiend as the gravitational force acting on a unit mass kept	a) It is given by $R_s = \frac{2GM}{c^2}$, where 'c' is velocity
at a point in the gravitational field.	of light, G universal gravitational constant and 'M'
b) If M is the mass of a body producing gravita-	the mass of the star that forms the black hole.
tional field and 'r' is the distance of the point in the	b) This equation implies that any star with mass 'M' can become a black hole if it can achieve
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will not be able to withstand, the core collapses Schwarzschild radius. and finally a black hole is formed. **STAGES IN THE BLACK HOLE FORMA-FRAME OF REFERENCE :** TION: i) A system of coordinate axes attached to an observer with the help of which the motion of an object can be described is called a frame of reference. Massive Star ii) The Cartesian system of coordinate axes can \downarrow (Energy source i.e, hydrogen decreases) be considered as the simplest frame of reference. Red giant stage In this system the position of an object at any \downarrow (Material particles blown off) moment can be described by three coordinates White dwarf stage (x,y,z) along the three coordinate axes X, Y,Z. **INERTIAL FRAMES OF REFERENCE :** \downarrow (Chandra Sekhar limit exceed) • Frames that move with constant velocity with Super Nova respect to each other are called "inertial frames of \downarrow (Further core collapse) reference". Neutron Star • An inertial reference frame is one in which \downarrow (Further core collapse) Newton's first law (law of inertia) is valid i.e 'no Black hole force no acceleration" is true. • When two observers from their respective inertial frames are watching an object, they report • Black hole formation is the last stage of the life i) different positions (s) ii) different velocities (v) cycle of a massive star. iii) same acceleration (a) • When the energy source i.e, hydrogen avail-• In Newtonian mechanics, invariant quantities are ability decreases, its core starts to shrink because mass, force, acceleration and time. For all observof gravity resulting in the formation of a Red giant. ers in inertial frames these will have same value. • At Red giant stage diameter becomes many more • The laws of physics are same in all inertial frames times the diameter of the original massive star due of reference. to nuclear fusion. f) A frame of reference attached to the earth may • After several million years, hydrogen available be considered as an inertial frame of reference, in the 'red giant' is exhausted completely. though strictly speaking it is a non-inertial frame. • The 'red giant' now releases material particles NON-INERTIAL FRAMES OF REFERpresent in it and forms a relatively dim star, known **ENCE**: Frames that move with acceleration are as 'white dwarf. called "non-inertial frames of reference". • At white dwarf state all the remaining material Ex. A lift carrying a person and moving up with acceleration. will be packed into a small volume, i.e., about one CONCEPTUAL QUESTIONS millionth the size of the original star, due to the Following physical quantity of a planet that strong gravitational pull. 1. revolves around Sun in an elliptical orbit is • When the mass of the white dwarf exceeds about constant. 1.4 times the mass of the Sun (Chandrasekhar's 1) Kinetic energy 2) Potential energy limit) it collapses further and the core temperature 3) Angular momentum 4) All rises over 100 billion degrees. 2. If the universal gravitational constant decreases • At this stage the repulsive force between mulei uniformly with time, then a satellite in orbit will still overcomes the attraction force of gravity giving maintain its 1) weight 2) tangential speed out an explosive shock wave, known as 'super-3) period of revolution 4) angular momentum nova'. 3. Two satellites of masses m_1 and m_2 ($m_1 > m_2$) are i) After 'supernova' stage, depending upon the revolving around earth in circular orbits of radii r, mass of the original star, pressure inside the 'left and $r_2(r_1 > r_2)$ respectively. Which of the following over' core becomes very large. statements is true regarding their velocities i) At this stage pressure electrons are forced to $V_1 \text{ and } V_2.$ 1) $V_1 = V_2$ 1) $V_1 = V_2$ 2) $V_1 < V_2$ 3) $V_1 > V_2$ 4) $\frac{V_1}{r_1} = \frac{V_2}{r_2}$ combine with protons reducing whole of the star into a dense ball of neutrons, known as 'Neutron Star'. • If mass of original star is very large, say 10 or more times the mass of the Sun, even neutrons

If the area swept by the line joining the sun and 14. the earth from Feb 1 to Feb 7 is 'A', then the area swept by the radius vector from Feb 7 to Feb 28 is 1)A 2) 2A 3) 3A 4) 4A 5. Two equal masses separated by a distance (d) attract each other with a force (F). If one unit of mass is transferred from one of them to the other, the force 15. 1) does not change 2) decreases by (G/d^2) 3) becomes d² times 4) increases by $(2G/d^2)$ If F_a and F_a are gravitational and electrostatic 6. forces between two electrons at a distance 0.1 m 16. then F_{g}/F_{s} is in the order of 1) 10⁴³ 3) 1035 2) 10^{-43} 4) 10^{-35} 7. Out of the following interactions the weakest is 1) gravitational 2) electromagnetic 3) nuclear 4) electrostatic 8. If the mean radius of earth is R, its angular velocity is ω and the acceleration due to gravity at the surface of the earth is 'g' then the cube of the radius 17. of the orbit of a satellite will be $1)\frac{Rg}{\omega^2} = 2)\frac{R^2g}{\omega} = 3)\frac{R^2g}{\omega^2} = 4)\frac{R^2\omega}{g}$ 9. The gravitational field is a conservative field. The work done in this field by moving an object from one point to another 18. 1) depends on the end-points only 2) depends on the path along which the object is moved 3) depends on the end-points as well as the path between the points. 4) is not zero when the object is brought back to 19. its initial position. 10. The tidal waves in the sea are primarily due to 1) the gravitational effect of the moon on the earth 2) the gravitational effect of the sun on the earth 3) the gravitational effect of the venus on the earth 4) the atmospheric effect of the earth itself 11. A hole is drilled through the earth along a diameter 20. and a stone is dropped into it. When the stone is at the centre of the earth, it has finite 1) weight 2) acceleration 3) P.E. 4) mass 12. If R=radius of the earth and g=acceleration due to gravity on the surface of the earth, the acceleration due to gravity at a distance r (< R)21. from the centre of the earth is proportional to 3) r^{-2} 2) r^2 4) r^{-1} 1) r 13. If R=radius of the earth and g=acceleration due to gravity on the surface of the earth, the acceleration due to gravity at a distance r(>R) from the centre of the earth is proportional to 2) r^2 1) r 3) r^{-2} 4) r^{-1}

For a satellite projected from the earth's surface with a velocity greater than orbital velocity the nature of the path it takes when its energy is negative, zero and positive respectively is

- 1) Elliptical, parabolic and hyperbolic
- 2) Hyperbolic, parabolic and elliptical
- 3) Elliptical, circular and parabolic
- 4) Parabolic, circular and Elliptical
- The period of a satellite moving in circular orbit near the surface of a planet is independent of 1) mass of the planet 2) radius of the planet
 - 3) mass of the satellite 4) density of planet
- Out of the following statements, the one which correctly describes a satellite orbiting about the earth is
 - 1) There is no force acting on the satellite
 - 2) The acceleration and velocity of the satellite are roughly in the same direction
 - 3) The satellite is always accelerating about the earth
 - 4) The satellite must fall, back to earth when its fuel is exhausted.
- When an astronaut goes out of his space-ship into the space he will
 - 1) Fall freely on the earth
 - 2) Go upwards
 - 3) Continue to move along with the satellite in the same orbit.
 - 4) Go spiral to the earth
- A body is dropped from a height equal to radius of the earth. The velocity acquired by it before touching the ground is

1) $V = \sqrt{2gR}$ 2) V = gR 3) $V = \sqrt{gR}$ 4) V = 2gR

- The earth retains its atmosphere. This is due to 1) The special shape of the earth
 - 2) The escape velocity being greater than the mean speed of the molecules of the atmospheric gases.
 - 3) The escape velocity being smaller than the mean speed of the molecules of the atmospheric gases. 4) The sun's gravitational effect.
- Ratio of the radius of a planet A to that of planet B is 'r'. The ratio of accelerations due to gravity for the two planets is x. The ratio of the escape velocities from the two planets is

1)
$$\sqrt{rx}$$
 2) $\sqrt{r/x}$ 3) \sqrt{r} 4) $\sqrt{x/r}$

- For a planet revolving round the sun, when it is nearest to the sun is
 - 1) K.E. is min and P.E. is max.
 - 2) Both K.E. and P.E. are min
 - 3) K.E. is max. and P.E. is min
 - 4) K.E. and P.E. are equal

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22.	When the height of a satellite increases from the	31.	A synchronous satelli	te should be at a proper
	surface of the earth.		height moving	
	1) PE decreases, KE increases		1) From West to East	in equatorial plane
	2) PE decreases, KE decreases		2) From South to Nor	th in polar plane
	3) PE increases, KE decreases		3) From East to West	in equatorial plane
	4) PE increases, KE increases		4) From North to Sou	th in polar plane
23.	When a satellite going round the earth in a	32.		velocity vector of a
	circular orbit of radius r and speed v loses some			nd the spin angular velocity
	of its energy, then r and v change as		vector of the earth are	
	1) r decreases, v increases		1) always in the same	direction
	2) both decrease 3) both increase		2) always in opposite	
	4) r increases, v decreases		3) always mutually per	
24.	If a satellite is moved from one stable circular orbit		4) inclined at 23 $1/2^{\circ}$	-
	to a farther stable circular orbit, then the following	33.	· ·	ore than 5 times the solar
	quantity increases	55.	the mass end their lives	
	1) Gravitational force 2) Gravitational P.E.		1) White dwarfs	2) Red giants
	3) linear orbital speed		3) Black dwarfs	ý Q
0.5	4) Centripetal acceleration	34.	· ·	ich one has the core of
25.	The energy required to remove an earth satellite	54.	highest density?	ich one has the core of
	of mass 'm' from its orbit of radius 'r' to infinity is		1) Neutron star	2) White dwarf
	1) $\frac{GMm}{r}$ 2) $\frac{-GMm}{2r}$ 3) $\frac{GMm}{2r}$ 4) $\frac{Mm}{2r}$		3) Yellow star	· · · · · · · · · · · · · · · · · · ·
	1) $\frac{1}{r}$ 2) $\frac{1}{2r}$ 3) $\frac{1}{2r}$ 4) $\frac{1}{2r}$		<i>,</i>	4) Red giant
26.	Assume that a satellite is revolving around earth in	35.	The radius of a bla	ck hole (R_B) and its
	a circular orbit almost close to the surface of earth.		0 1 1 1 1 1	(\mathbf{D}) 1.1
	The time period of revolution of satellite is (Radius		Schwartzchild radius	(R_S) are related as
	of earth is $6400 \text{ km}, \text{g} = 9.8 \text{ ms}^{-2}$)		1) $R_{B} > R_{S}$	2) $R_B \ge R_S$
	1) 5068 sec2) 5068 min 3) 24 hour 4) 1 year			
27.	A body has weight (w) on the ground. The work		3) $R_B \leq R_S$	4) $R_B = R_S = \text{Infinity}$
	which must be done to lift it to a height equal to	36.	A black hole has	
	the radius of the earth is		1) zero volume and zer	to density
	1) equal to WR2) greater than WR		2) zero density and infi	
	3) less than WR 4) we can't say		3) zero volume and infi	nite density
28.	Moon is revolving in a circular orbit with a		4) infinite volume and in	nfinite density
	uniform velocity V_0 . If the gravitational force	37.	The boundary of a blac	ck hole is called
	suddenly disappears, the moon will		1) event horizon	2) Schwartzchild radius
	1) continue to move in the same orbit		3) Chandra sekhar limi	t 4) Einstein's space time
	2) move with a velocity V_0 tangentially to the orbit	38.	If M is mass of the sun	, then the mass of a white
20	3) fall down freely 4) ultimately comes to rest		dwarf star may be	
29.	The time period of revolution of geostationary		1) M 2) 2M	3) 3M 4) 4M
	satellite with respect to earth is	39.	Chandra Sekhar limit i	s about
30.	1) 24 hrs 2) 1 year 3) Infinity 4) Zero A relay satellite transmits the television programme		1) 2.4 times the solar n	nass
50.	from one part of the world to another part		2) 1.4 times the solar n	nass
	continuously because its period		3) 14 times the solar m	
	1) is greater than the period of the earth about its axis		4) 24 times the solar m	
	2) is less than period of rotation of the earth	40.	During the transformat	ion of a massive star ulti-
	about its axis.		÷	e, which of the following
	3) has no relation with the period of rotation of		sequence is correct?	
	the earth about its axis.		/	ova stage, white dwarf stage
	4) is equal to the period of rotation of the earth		<i>,</i>	giant stage, supernova stage
	about its axis.		, U I	ernova stage, red giant stage
			,	te dwarf stage, supernova
			stage	
L		1		

height moving			
4 N TH			
1) From West to East	t in equatorial plane		
2) From South to North in polar plane			
3) From East to West in equatorial plane			
	4) From North to South in polar plane		
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1) always in the same			
2) always in opposite			
3) always mutually pe	-		
4) inclined at 23 $1/2^{\circ}$			
	nore than 5 times the solar		
the mass end their live	es as		
1) White dwarfs	2) Red giants		
3) Black dwarfs	4) Black holes		
· · · · · · · · · · · · · · · · · · ·	hich one has the core of		
highest density?			
1) Neutron star	2) White dwarf		
3) Yellow star	4) Red giant		
,	, C		
The radius of a bl	ack hole (R_B) and its		
Schwartzchild radius	(R_S) are related as		
1) $R_B > R_S$	$2) R_{B} \geq R_{S}$		
3) $R_B \leq R_S$	4) $R_B = R_S = \text{Infinity}$		
A black hole has			
A black hole has 1) zero volume and ze	ero density		
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1) zero volume and ze	finite volume		
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 zero volume and zero zero density and inf zero volume and inf infinite volume and inf infinite volume and infinite volum	Finite volume finite density infinite density ack hole is called 2) Schwartzchild radius hit 4) Einstein's space time h, then the mass of a white		
 zero volume and zero zero density and inf zero volume and inf zero volume and inf infinite vol	Einite volume finite density infinite density ack hole is called 2) Schwartzchild radius hit 4) Einstein's space time h, then the mass of a white 3) 3M 4) 4M		
 zero volume and zero 2) zero density and inf zero volume and inf zero volume and inf infinite volume and infinite volu	Finite volume finite density infinite density ack hole is called 2) Schwartzchild radius hit 4) Einstein's space time h, then the mass of a white 3) 3M 4) 4M is about		
 zero volume and zero 2) zero density and inf zero volume and inf zero volume and inf infinite volume and infinite vol	Finite volume finite density infinite density ack hole is called 2) Schwartzchild radius hit 4) Einstein's space time h, then the mass of a white 3) 3M 4) 4M is about mass		
 zero volume and zero 2) zero density and inf zero volume and inf zero volume and inf infinite volume and infinite infinite volume and infinite	Einite volume finite density infinite density ack hole is called 2) Schwartzchild radius hit 4) Einstein's space time h, then the mass of a white 3) 3M 4) 4M is about mass mass		
 zero volume and zero 2) zero density and inf zero volume and inf zero volume and inf infinite volume and infinite vol	Einite volume finite density infinite density ack hole is called 2) Schwartzchild radius hit 4) Einstein's space time h, then the mass of a white 3) 3M 4) 4M is about mass mass mass		
 zero volume and zero 2) zero density and inf zero volume and inf zero volume and inf infinite volume and infinite volume	Finite volume finite density infinite density ack hole is called 2) Schwartzchild radius hit 4) Einstein's space time h, then the mass of a white 3) 3M 4) 4M is about mass mass mass mass mass		
 zero volume and zero 2) zero density and inf zero volume and inf zero volume and inf zero volume and inf infinite volume and volume and volume and volume and volume and	Einite volume finite density infinite density ack hole is called 2) Schwartzchild radius hit 4) Einstein's space time h, then the mass of a white 3) 3M 4) 4M is about mass mass mass mass ation of a massive star ulti-		
 zero volume and zero 2) zero density and inf zero volume and inf zero volume and inf infinite volume and infinite infinite volume and	Finite volume finite density infinite density ack hole is called 2) Schwartzchild radius hit 4) Einstein's space time h, then the mass of a white 3) 3M 4) 4M is about mass mass mass mass mass		
 zero volume and zero 2) zero density and inf zero volume and inf zero volume and inf infinite volume and infinite infinite volume and	Finite volume finite density infinite density ack hole is called 2) Schwartzchild radius nit 4) Einstein's space time n, then the mass of a white 3) 3M 4) 4M is about mass mass mass mass ation of a massive star ulti- le, which of the following		
 zero volume and zero 2) zero density and inf zero volume and inf zero volume and inf infinite volume and sekhar limit infinite volume a	Einite volume finite density infinite density ack hole is called 2) Schwartzchild radius hit 4) Einstein's space time h, then the mass of a white 3) 3M 4) 4M is about mass mass mass ation of a massive star ulti- ble, which of the following nova stage, white dwarf stage		
 zero volume and zero 2) zero density and inf zero volume and inf zero volume and inf infinite volume and sekhar limit infinite volume a	Finite volume finite density infinite density ack hole is called 2) Schwartzchild radius nit 4) Einstein's space time n, then the mass of a white 3) 3M 4) 4M is about mass mass mass mass ation of a massive star ulti- le, which of the following		
 zero volume and zero 2) zero density and inf zero volume and im zero volume and im infinite volume and im infinite volume and im infinite volume and im event horizon Chandra sekhar lim If M is mass of the surdwarf star may be M 2) 2M Chandra Sekhar limit 2.4 times the solar 14 times the solar 14 times the solar 24 times the solar 16 a black horizon Sequence is correct? Red giant stage, super White dwarf stage, red 	Einite volume finite density infinite density ack hole is called 2) Schwartzchild radius hit 4) Einstein's space time h, then the mass of a white 3) 3M 4) 4M is about mass mass mass ation of a massive star ulti- ble, which of the following nova stage, white dwarf stage		
 zero volume and zero 2) zero density and inf zero volume and inf zero volume and inf infinite volume ask of the sum of the solar infinite volume and the solar infinite volume and	Einite volume finite density infinite density ack hole is called 2) Schwartzchild radius nit 4) Einstein's space time n, then the mass of a white 3) 3M 4) 4M is about mass mass mass mass ation of a massive star ulti- ble, which of the following nova stage, white dwarf stage d giant stage, supernova stage		
 zero volume and zero 2) zero density and inf zero volume and inf zero volume and inf infinite volume ask of the sum of the solar infinite volume and the solar infinite volume and	Einite volume finite density infinite density ack hole is called 2) Schwartzchild radius nit 4) Einstein's space time n, then the mass of a white 3) 3M 4) 4M is about mass mass mass mass ation of a massive star ulti- ble, which of the following nova stage, white dwarf stage d giant stage, supernova stage pemova stage, red giant stage		

41.		called fictitious force such as	ORBI	TALVELOCITY - ESCAPE VELOCITY :
	centrifugal force ari	ses only in	5. A	satellite of mass 'm' revolves round the earth o
	1) Inertial frames	-	ll m	ass 'M' in a circular orbit of radius 'r' with a
	2)Non-intertial fran	nes		ngular velocity ' ω '. If the angular velocity is
	3) Both inertial and			$\sqrt{8}$ the radius of the orbit will be
	4) Rigid frames			
42.		the gravitational field) 4r 2) 2r 3) 8r 4) r
72.	propagates is	the gravitational new		particle falls towards earth from infinity. Th
	1 1 0	d officiet in the supervise	Ve Ve	elocity with which it reaches earth's surface is.
	/ 1 1	ed of light in vacuum	1)) v = 2gR 2) v = $\sqrt{2gR}$
	· ·	ed of light in vacuum		
	ý 1	eed of light in vacuum	3)) $v = \sqrt{gR}$ 4) $v = R/g$
	4) Either less or m	ore than the speed of light in		he escape velocity on a planet is 'v'. If the radiu
	vacuum			f the planet contracts to 1/4th of present valu
	K	<u>XEY</u>		
				ithout any change in its mass, the escape velocit ill be
	1.3 2.4	3.2 4.3 5.2	11	
	6.2 7.1	8.3 9.1 10.1	/	halved 2) doubled
	11.4 12.1	13.3 14.1 15.3	<i>′</i>) quadrupled 4) becomes one fourth
	16.3 17.3	18.3 19.2 20.1		he escape velocity from the earth for a rocket
	21.3 22.3	23.1 24.2 25.3		1.2 km/s ignoring air resistance. The escape velocit
	26.1 27.1	28.2 29.3 30.4	of	f 10 mg grain of sand from the earth will be
	31.1 32.1	33.4 34.1 35.3	1)) 0.112 km/s 2) 11.2 km/s
	36.3 37.1	38.1 39.2 40.4	3) 1.12 km/s 4) 0.0112 kms^{-1}
	41.3 42.1	0012 1011	11 ⁽	he escape velocity from the surface of the eart
		/EL - I		fradius R and density ρ
		IVERSAL LAW OF		•
		ITATION :		$2R\sqrt{\frac{2\pi\rho G}{3}} \qquad 2) \ 2\sqrt{\frac{2\pi\rho G}{3}}$
1				$2R\sqrt{\frac{2}{3}}$ 2) $2\sqrt{\frac{3}{3}}$
1.	6	force between two bodies is		, - , -
		the distance between their		$2 R \sqrt{R}$
		the mass of first body is 800	3)) $2\pi \sqrt{\frac{R}{g}}$ 4) $\sqrt{\frac{2\pi G\rho}{R^2}}$
	kg, then the mass of	-		18
	1) 1000 kg	2) 1250 kg 3) 1500 kg		body is projected vertically up from surface of
_	4) 2000kg			e earth with a velocity half of escape velocity
2.	•	kg mass are placed on x and		he ratio of its maximum height of ascent an
	•	of 1 metre from the origin and	ra	idius of earth is
		aced at the origin. Then the	1)	$(1:1 \ 2) \ 1:2 \ 3) \ 1:3 \ 4) \ 1:4$
	U U	nal force on 1 kg mass is	11. Tv	wo satellites are revolving round the earth i
	1) 7G 2) G	3) 5G 4) 3G	ci	rcular orbits of radii in the ratio 1:2. Their orbita
3.		res of same material and	l ve	elocities are in the ratio of
	radius 'r' are in co	ontact with each other. The	1	$(1:2 2) \sqrt{2} : 1 3) \ 2\sqrt{2} : 1 4) \ 8 : 1$
	gravitational force	e of attraction between the	11 *	
	spheres is given by			the mass of earth were 4 times the present mass
		$K / r^{3} 3$) F=K / 4r ² 4) Kr ²		e mass of the moon were half the present mas
4.	· · · ·	qual mass move in a circle of		nd the moon were revolving round the earth a
	-	he action of their mutual		e same present distance, the time period of
		on. If the mass of each particle	re	evolution of the moon would be
	-	_	1)) 56 days 2) 28 days 3) 14 days 4) 7 days
	is m, the speed of e	ach particle is		ENERGY OF A SATELLITE
	Gm	Gm Gm 2Gm	13. T	wo satellites of masses 400 kg, 500 kg ar
I	1) $\sqrt{\frac{3}{2}}$ 2) $\sqrt{\frac{3}{2}}$	$\frac{Gm}{2r}$ 3) $\sqrt{\frac{Gm}{4r}}$ 4) $\sqrt{\frac{2Gm}{r}}$		evolving around earth in different circular orbit
			10 10	working around cardinin anterent encardi oron
	vr	$2r$ $\sqrt{4r}$ \sqrt{r}		f radii r_1 , r_2 such that their kinetic energies an

14.	The K.E. of a satellite is 10 ⁴ J, its P.E. is	24.	Two stars have masses $5 \times 10^{30} kg$ and
	1) -10^4 J 3) -2×10^4 J 4) -4×10^4 J		$7.5 \times 10^{30} kg$ respectively. If they ultimately con-
15.	A satellite moves around the earth in a circular		vert into black holes, the ratio of Schwartzschild
13.	orbit with speed 'V'. If 'm' is mass of the satellite		ratio of the black holes is
	then its total energy is		1)2:3 2)4:9 3)3:2 4)9:4
			KEY
	1) $\frac{1}{2}$ mv ² 2) mv ² 3) $-\frac{1}{2}$ mv ² 4) $\frac{3}{2}$ mv ²		1.2 2.3 3.1 4.1 5.1
10			6.2 7.2 8.2 9.1 10.3
16.	If R is radius of the earth and W is work done in		11.3 12.3 13.1 14.3 15.3
	lifting a body from the ground to an altitude R,		16.1 17.2 18.4 19.3 20.3
	the work which should be done in lifting it further to twice that altitude is		21.4 22.2 23.2 24.1
	1) $W/2$ 2) W 3) $W/3$ 4) $3W$		<u>LEVEL - II</u>
	GEOSTATIONARY SATELLITE		NEWTON'S LAW OF GRAVITATION
17.	A satellite is orbiting round the earth. If both	25.	If the mass of one particle is increased by 50 %
- / -	gravitational force and centripetal force on the		and the mass of another particle is decreased by
	satellite is 'F' then net force acting on the satellite		50%, the force between them
	to revolve round the earth is		1) decreases by 25% 2) decreases by 75%
	1) F/2 2) F 3) 2F 4) Zero	26	3) increases by 25% 4) does not change The gravitational force between two bodies is
18.	The minimum number of geostationary satellites	26.	decreased by 36 % when the distance between
	required to televise a programme all over the		them is increased by 3m. The initial distance
	earth is		between them is
	1) 2 2) 6 3) 4 4) 3 BLACK HOLES		1) 6 m 2) 9 m 3) 12 m 4) 15 m
19.	If two stars of masses in the ratio 2 : 3 become	27.	If the distance between two bodies is increased by
19.	black holes, their radii will be in the ratio of		25%, then the % change in the gravitational force is
	1)4:9 2) $3:2$ 3) $2:3$ 4) $9:4$		1) Decreases by 36% 2) Increases by 36%
20			3) Increases by 64% 4) Decreases by 64%
20.	When a star of mass $9 \times 10^{30} kg$ ends as a black	28.	Masses 2 kg and 8 kg are 18 cm apart. The point
	hole, the Schwartzschild radius of the star is		where the gravitational field due to them is zero is 1×6 are from 2×10^{-10}
	$\left(G = 6.7 \times 10^{-11} Nm^2 kg^{-2}\right)$		 6 cm from 8 kg mass 6 cm from 2 kg mass
	1) 13.4 m 2) 6.7 m 3) 13.4 km 4) 26.8 km		3) 1.8 cm from 8 kg mass
21.	Two lead spheres of same radius are in contact		4) 9 cm from each mass
	with each other. The gravitational force of at-	29.	Mass of the earth is 81 times that of the moon. If
	traction between them is F. If two lead spheres		the distance between the centre of the earth and
	of double the previous radius are in contact with		the center of moon is d then the distance from the
	each other, the gravitational force of attraction		centre of the earth at which gravitational field
	between them will be		strength due to earth - moon system is zero is
22	1) 2F 2) 32F 3) 8F 4) 16F		1) d/81 2) 9d/10 3) d/10 4) 8d/9
22.	Two particles of masses 'm' and '2m' are at a distance '2r' apart at the ande of a straight line	30.	Two lead balls of masses m and 5m having radii R and 2P are compared by 12P. If they attract each
	distance '3r' apart at the ends of a straight line AB. C is the centre of mass of the system. The		and 2R are separated by 12R. If they attract each other by gravitational force, the distance covered
	magnitude of the gravitational intensity due to the		by small sphere before they touch each other is
	magnitude of the gravitational intensity due to the		1) 10 R 2) 7.5 R 3) 9 R 4) 2.5 R
	masses at C is $7Gm$ 1) Zero 2) $\frac{7Gm}{4r^2}$ 3) $\frac{9Gm}{4r^2}$ 4) $\frac{3Gm}{2r^2}$	OR	BITALVELOCITY & ESCAPE VELOCITY
	$4r^2$ $4r^2$ $2r^2$	31.	A satellite is revolving around earth in a circular
23.	Two masses 'M' and '4M' are at a distance 'r'		orbit of radius equal to diameter of earth. The
	apart on the line joining them, 'P' is point where		minimum % increase in the speed of that satellite
	the resultant gravitational intesity is zero (such a point called null point). The distance of 'P' from		so that it escapes from earth's gravity is
			1) 100 % 2) 82.8 % 3) 50 % 4) 41.4 %
1	the mass 'M' is. 1) $\frac{r}{5}$ 2) $\frac{r}{3}$ 3) $\frac{2r}{3}$ 4) $\frac{4r}{5}$		
	-, 5 -, 3 -, 3 , 5		
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- 32. The speed of a satellite that revolves around earth at a height 3R from earth's surface is ($g = 10 \text{ m/s}^2$ at the surface of earth, radius of earth R = 6400 km.)
 - 1) $2\sqrt{2}$ kms⁻¹ 2) 4 kms⁻¹

3) $4\sqrt{2}$ kms⁻¹ 4) 8 kms⁻¹

- 33. Two satellites P, Q are revolving around earth in different circular orbits. The velocity of P is twice the velocity of Q. If the height of P from earth's surface is 1600 km. The radius of orbit of Q is (radius of earth R = 6400 km). 2) 20000 km
 - 1) 1600 km
 - 3) 32000 km 4) 40000 km
- 34. Two satellites M and N go around the earth in circular orbits at heights of R_M and R_N respectively from the surface of the earth. Assuming the earth to be a uniform sphere of radius R_{μ} , the ratio of

the velocities of the satellites $\frac{V_M}{V_M}$ is

1)
$$\left(\frac{R_M}{R_N}\right)^2$$

2) $\sqrt{\frac{R_N + R_E}{R_M + R_E}}$
3) $\frac{R_N + R_E}{R_M + R_E}$
4) $\sqrt{\frac{R_N}{R_M}}$

35. The escape velocity from an altitude equal to radius of earth above earth's surface is (escape velocity from surface of earth is 11.2kms⁻¹) $1s^{-1}$

1)
$$5.6 \text{ kms}^{-1}$$
 2) 7.92 km

- 3) 2.8 kms⁻¹ 4) 11.2 kms⁻¹
- 36. A particle is kept at rest at a distance R (Earth's radius) above the earth's surface. The minimum speed with which it should be projected so that it does not return is

1)
$$\sqrt{\frac{GM}{R}}$$
 2) $\sqrt{\frac{GM}{2R}}$ 3) $\sqrt{\frac{GM}{3R}}$ 4) $\sqrt{\frac{GM}{4R}}$

37. If the radius of the earth is reduced by 1 % keeping the mass constant. The escape velocity will

1) increase by 0.5%2) decrease by 0.5%3) decrease by 11% 4) remain same

38. The radius of the earth is R and acceleration due to gravity at its surface is g. The minimum speed with which a body must be thrown from the surface of the earth so as to reach a height of R/4 is

1)
$$\sqrt{gR}$$
 2) $\sqrt{\frac{gR}{2}}$ 3) $\sqrt{\frac{gR}{5}}$ 4) $\sqrt{\frac{2gR}{5}}$

- 39. A space-ship is launched into a circular orbit close to the surface of the earth. The additional velocity now to be imparted to the space-ship in the orbit to overcome the gravitational pull is 1) 11.2 km/sec 2) 8 km/sec 4) $1.414 \times 8 \text{ km/sec}$ 3) 3.2 km/sec40. The moon escapes for ever, if the minimum increase in its velocity is 1) 200 % 2) 41.4 % 3) 50 % 4) 100 % **ENERGY OF A SATELLITE AND GEOSTATIONARY SATELLITE** The gravitational P.E. of a rocket of mass 100 kg 41. at a distance of 10^7 m from the earths centre is -4×10^9 J. The weight of the rocket at a distance of 10^9 m from the centre of the earth is 1) 4 x 10⁻² N 2) 4 x 10⁻⁹ N 3) 4 x 10⁻⁶ N 4) 4 x 10⁻³ N 42. A man weighs 75 kg on the surface of the earth. His weight in a geostationary satellite is 2) 150 kg 3) zero 4) 75/2 kg 1) infinity
- Two identical particles each of mass 'm' start 43. moving towards each other from rest from infinite separation under gravitational attraction. Their relative velocity of approach at separation 'r'is

1)
$$\sqrt{\frac{Gm}{r}}$$
 2) $\sqrt{\frac{2Gm}{r}}$ 3) $2\sqrt{\frac{Gm}{r}}$ 4) $\sqrt{\frac{Gm}{2r}}$

44. Three identical particles each of mass "m" are arranged at the corners of an equilateral triangle of side "L". If they are to be in equilibrium, the speed with which they must revolve under the influence of one another's gravity in a circular orbit circumscribing the triangle is

1)
$$\sqrt{\frac{3Gm}{L}}$$
 2) $\sqrt{\frac{Gm}{L}}$ 3) $\sqrt{\frac{Gm}{3L}}$ 4) $\sqrt{\frac{3Gm}{L^2}}$

45. A small body is initially at a distance 'r' from the centre of earth. 'r' is greater than the radius of the earth. If it takes W joule of work to move the body from this position to another position at a distance 2r measured from the centre of earth, how many joules would be required to move it from this position to a new position at a distance of 3r from the centre of the earth.

The mass of the sun is approximately $2x10^{30}$ kg. 46. The Schwarzschild radius for the mass of a star that is ten times the mass of sun is nearly 1) 3km 2) 30 km 3) 300 km 4) 0.3 km

KEY25. 126.327.128.229.235.236.137.128.429.235.236.137.138.439.340.241.142.343.344.2HINTSEscape velocity from the surface of earth is 11.2Error evelocity from the surface of earth is 11.225.
$$F_1 = G\frac{m,m_3}{d^2}$$
41.12.526. $F_{\frac{1}{2}} = \frac{d_1^2}{d_1^2}$ 71.12.827. $F_1 = G\frac{(m_1 + \frac{m_1}{2})(m_2 - \frac{m_2}{2})}{d^2}$ 31.72.828. $\frac{M}{F_2} = \frac{d_1^2}{d_1^2}$ 11.21.428. $\frac{m_1}{T_2} = \frac{d_1^2}{d_1^2}$ 11.14.129.19.4km/s2.11.929.19.4km/s2.11.929.19.4km/s2.11.928. $\frac{m_1}{m_1} = \frac{1}{100} F_1; d_2 = (d_1 + 3)$ 1.91.928. $\frac{m_1}{m_1} = \frac{m_2}{m_1} |(GR)$ 1.91.91.929.19.23.03.03.0 $\frac{Mm}{6R} = 3$ 29.19.22.12.23.1 $\frac{m_1}{3}$ 30.2.11.41.91.91.41.928. $\frac{Mm}{m_1} = \frac{m_2}{100} |(GR)$ 1.91.91.431.2.11.91.41.41.932.1.91.41.41.91.233.2.42.41.91.41.430. $\sqrt{2}(4R) = \sqrt{\frac{K}{16} (4R)$ 1.91.931. $\sqrt{R} = 2$

- 57. The altitude of geostationary satellite is nearly 6 times the radius of the earth. The period of revolution of an identical satellite revolving at an altitude 0.75 times the radius of the earth will be
- 1) 4 hrs 2) 3 hrs 3) 12 hrs 4) 2 hrs 58. A satellite is geostationary in a particular orbit. It is allowed to go to another orbit having orbital radius 2 times that of the earlier orbit from the centre of the earth. The time period in the second orbit is 1) 48hrs 2) 24hrs 3)48 $\sqrt{2}$ hrs 4) 24 $\sqrt{2}$ hrs
- 59. Explorer- 38, a radio-activity research satellite of mass 200 kg circles the earth in an orbit of radius 3R/2, where R is the radius of the earth. Assuming the gravitational pull on a mass of 1 kg at the earth's surface to be 10 N, the pull on the satellite is 1) 889 N 2) 4500 N 3) 9000 N 4) None
- 60. A geo-stationary satellite is orbitting the earth at a height 6R above the surface of the earth, where R is the radius of earth. The time period of another satellite revolving around earth at a height 2.5 R from earth's surface is
- 1) 12√2 Hr. b 2) 12 hr 3) 6√2 hr 4) 6 hr
 61. Particles each of mass M are placed along x-axis at x=1m, x=2m, x=4m, x=8m,.... etc to infinity. Gravitational field strength at the origin due to this system of particles is
- 2GM 2) 2GM/3 3) 4GM/3 4) 5GM/4
 Three particles, each of mass 'm' are situated at the vertices of an equilateral triangle of side 'a'. The only forces acting on the particles are their mutual gravitational forces. It is desired that each particle should move in a circle while maintaining the original mutual separation 'a'. Then their time period of revolution is

1)
$$2\pi\sqrt{\frac{a^2}{3Gm}}$$

2) $2\pi\sqrt{\frac{a^3}{3Gm}}$
3) $2\pi\sqrt{\frac{3a^4}{Gm}}$
4) $2\pi\sqrt{\frac{a^4}{Gm}}$

63. If d is the distance between the centres of the earth of mass M_1 and moon of mass M_2 , then the velocity with which a body should be projected from the mid point of the line joining the earth and the moon, so that it just escapes is

1)
$$\sqrt{\frac{G(M_1 + M_2)}{d}}$$
 2) $\sqrt{\frac{G(M_1 + M_2)}{2d}}$
3) $\sqrt{\frac{2G(M_1 + M_2)}{d}}$ 4) $\sqrt{\frac{4G(M_1 + M_2)}{d}}$

64. Particles of masses m_1 and m_2 are at a fixed distance apart. If the gravitational field strength at m_1 and m_2 are l_1 and l_2 respectively. Then,

1)
$$m_1l_1 + m_2l_2 = 0$$

2) $m_1l_2 + m_2l_1 = 0$
3) $m_1l_1 - m_2l_2 = 0$
4) $m_1l_2 - m_2l_1 = 0$

<u>KEY</u>				
47.4	48.1	49.1	50.3	51.1
52.4	53.1	54.1	55.1	56.3
57.2	58.3	59.1	60.3	61.3
62.2	63.4	64.1		

HINTS

48. From $g = \frac{4}{3}\pi\rho GR, \frac{g_1}{g_2} = \frac{1}{2}$

from
$$\frac{V_1}{V_2} = \sqrt{\frac{2gR}{2(2g)(2R)}}$$

$$V_2 = 2V_1 = 22kms^{-1}$$
49. Velocity at infinity

$$V_{\infty} = \sqrt{V^2 - V_e^2} = \sqrt{\left[2(11.2)\right]^2 - (11.2)^2}$$

2GM

R

50.
$$\sqrt{\frac{GM}{R+h}} = \frac{1}{2} \left\lfloor \sqrt{\frac{2}{3}} \right\rfloor$$

51.
$$\Delta PE = \frac{GMm}{R} \left[1 - \frac{1}{2} \right] = \frac{GMm}{2R}$$

But

$$gR = \frac{GM}{R}$$
$$\therefore \Delta PE = \frac{mgR}{2}$$

56.
$$\frac{1}{2}mv^{2} = \frac{(mg)Rh}{(R+h)} = \frac{(mg)R(nR)}{(R+nR)}$$

57.
$$\frac{T_1}{T_2} = \sqrt{\frac{R_1^3}{R_2^3}}$$

58. $\frac{T_1}{T_2} = \sqrt{\frac{R_1^3}{R_2^3}} = \sqrt{\frac{R^3}{(2R)^3}}$

59.
$$\frac{g_1}{g_2} = \frac{R_2^2}{R_1^2} = \frac{\left(\frac{3}{2}R\right)^2}{R^2} = \frac{9}{4}$$

$$\therefore g_2 = \frac{4g}{9}$$

Weight =
$$200 \left[\left(\frac{4}{9} \right) g \right] = \frac{8000}{9} = 889N$$

 $T_1 = \sqrt{R_1^3} = \sqrt{(7R)^3}$

60.
$$\frac{T_1}{T_2} = \sqrt{\frac{R_1}{R_2^3}} = \sqrt{\frac{(7R)}{\left(\frac{7R}{2}\right)^3}}$$

61.
$$E_G = GM \left[\frac{1}{1^2} + \frac{1}{2^2} + \frac{1}{4^2} + \dots \right]$$

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$$= GM\left[\frac{1}{1-r}\right] = \frac{4}{3}GM$$
62. $mr\omega^2 = \sqrt{3}\frac{Gm^2}{a^2}$ and $r = \frac{a}{\sqrt{3}}$
63. PE at mid point $= \frac{-2GMm}{d}(M_1 + M_2)$
 $\Rightarrow \frac{1}{2}mV_e^2 = \frac{2Gm}{d}(M_1 + M_2)$
64. $I_1 = \frac{Gm_2}{d^2}$ and $I_2 = \frac{Gm_1}{d^2}$
 $\therefore \frac{G}{d^2} = \frac{I_1}{m_2} = \frac{-I_2}{m_1}$
65. Match List I with List II
List - I List - II a) White Dwarf c) 1.4 solar masses
b) Northern star f) Mass of star must be greater than 10 solar masses
c) Black holes g) Mass of the star must be less than 10 solar masses
d) Chandra Sekhar h) Mass of remaining star is greater than 3 solar masses
1) $a - g, b - f, c - h, d - e$
2) $a - f, b - g, c - e, d - h$
3) $a - g, b - f, c - e, d - h$
4) $a - f, b - e, c - g, d - h$
66. Which of the following is true for a satellite in a circular orbit a) It is freely falling body b) Its speed is constant

a) It is freely falling body b) Its speed is constant
c) It suffers no acceleration
d) It does not require energy for motion in the orbit

ololi	
1) a,b and c	2) a,b and d
3) a and b	4) a,b,c and d

- 67. Inside a uniform spherical shell
 a) Gravitational potential is zero
 b) Gravitational field is zero
 c) Gravitational potential is constant
 d) Gravitational field is constant
 1) a and b 2) b and d 3) b and c 4) a and c
 68. When a planet moves around the sun :
 - When a planet moves around the sun : a) Its angular momentum remains constant
 - b) It moves faster when it is nearer to the sun

c) Its total energy increases when it goes nearer to the sun.

d) Its potential energy decreases when it goes nearer to the sun.

1) only a and b are true 2) only b and c are true 3) only a, b and d are true 4) all are true

69. The magnitudes of the gravitational field at distances r_1 and r_2 from the centre of a uniform sphere of radius R and mass M are F_1 and F_2 respectively. Then :

a)
$$\frac{F_1}{F_2} = \frac{r_1}{r_2}$$
 if $r_1 < R$ and $r_2 < R$
b) $\frac{F_1}{F_2} = \frac{r_2^2}{r_1^2}$ if $r_1 > R$ and $r_2 > R$
c) $\frac{F_1}{F_2} = \frac{r_1}{r_2}$ if $r_1 > R$ and $r_2 > R$
d) $\frac{F_1}{F_2} = \frac{r_1^2}{r_2^2}$ if $r_1 < R$ and $r_2 < R$

1) only a and b are true 2) only b and c are true 3) only a,b and d are true 4) all are true

70. A satellite is revolving round the earth in an elliptical obit :

a) Gravitational force exerted by earth is equal to centripetal force at some points only.

b) Power associated with gravitational force is zero at every point

c) Work done by gravitational force is zero in some small parts of the orbit.

d) At some points, magnitude of gravitational force is greater than that of centripetal force.

- 1) Only a and b are true
- 2) Only b and c are true
- 3) Only a, b and d are true
- 4) Only a,c and d are true
- 71. Let V, T, L, K and r denote the speed, time period, angular momentum, kinetic energy and radius of satellite in a circular orbit, their

a) Var⁻¹ b) Lar^{1/2} c) Tar^{3/2} d) Kar⁻²

1) Only a and b are true 2) Only b and c are true

- 3) Only a, b and d are true
- 4) Only a,b and care true

Assertion & Reason :

72. A): For a planet revolving around the sun its velocity is more when it is closest to the sun.P): Low of concentration of consultant provider the sun is a supervised of the supervised of the

R): Law of conservation of angular momentum is the underlying principle

1) Both A and R are true and R is correct explanation of A

2) Both A and R are true and R is not the correct explanation of A

- 3) A is true but R is false
- 4) A is false but R is true

 73. A) : A body is first taken along a smooth inclined plane through a certain height. In the next case it is lifted vertically upwards to same height, work done by gravity in both cases is same R) : Gravitational force is conservative force 1) Both A and R are true and R is correct explanation of A 2) Both A and R are true and R is not the correct explanation of A 3) A is true but R is false 4) A is false but R is true 	 The eacape velocity of a body on the earth's surface is V_E. A body is thrown up with a speed √5V_E. Assuming that the sun and planets do not influence the motion of the body, velocity of the body at infinite distance is (Eng-2004) 0 V_E √2V_E V_E The radius in kilometers, to which the present radius of the earth (R = 6400 km) is to be compressed so that the escape velocity is increased ten times is: (Med. 2003) 6.4 644 640 4800
 74. A): When a body is dropped into a tunnel dug along the diameter of the earth it executes simple harmonic motion. R): At any position, magnitude of Gravitational force experienced by the body is directly propotional to the displacement of the body as measured from center of the earth. 1) Both A and R are true and R is correct explanation of A 2) Both A and R are true and R is not the correct explanation of A 3) A is true but R is false 4) A is false but R is true 75. A): A satellite revolving round the earth with uniform speed moves tangentially to the orbit with same speed when gravity suddenly disappears. 	 A satellite is launched into a circular orbit of radius "R' around the earth while a second satellite is launched into an orbit of radius 1.02 R. The percentage difference in the time periods of the two satellites is : (Eng-2003) 1) 0.7 2) 1.0 3) 1.5 4) 8 Two satellites S₁ and S₂ are revolving round a planet in coplanar and concentric circular orbits of radii R₁ and R₂ in the same direction respectively. Their respective periods of revolution are 1 hr. and 8 hr. The radius of the orbit of satellite S₁ is equal to 10⁴ km. Their relative speed when they are closest, in kmph is : (Med-2002) 1) π×10⁴ 2) π×10⁴ 3) 2π×10⁴ 4) 4π×10⁴
R) : This is due to inertia of motion 1) Both A and R are true and R is correct expla- nation of A 2) Both A and R are true and R is not the correct explanation of A 3) A is true but R is false 4) A is false but R is true $\frac{KEY}{65.1 66.2 67.3 68.3 69.1}{70.4 71.2 72.1 73.1 74.1}$ 75.2 $\frac{PREVIOUS EAMCET QUESTIONS}{ENGINEERING \& MEDICAL}$ 1. A spaceship is launched into a circular orbit of radius 'R' close to the surface of earth. The addi- tional velocity to be imparted to the spaceship in the orbit to overcome the earth's gravitational pull is : (g = acceleration due to gravity) $\frac{(Med-2004)}{1) 1.414Rg} = \frac{2}{4\pi \times 10^4}$	 A body is projected up with a velocity equal to 3/ 4 th of the escape velocity from the surface of the earth. The height it reaches is (Radius of the earth is R) 2002E 1) 10R/9 2) 9R/7 3) 9R/8 4) 10R/3 2001E Mass M is divided into two parts Xm and (1– X)m. For a given seperation the value of X for which the gravitational attraction between the two pieces becomes maximum is 1) 1/2 2) 3/5 3) 1 4) 2 2001M When a satellite going around the earth in a circular orbit of radius r and speed v loses some of its energy, then r and v both increase 2) r and v both decrease r will increase and v will decrease r will decrease and v will increase Two satellites of masses 50 kgs and 100 kgs revolve around the earth in circular orbit of radii 9R and 16 R respectively, where 'R' is the radius of the earth. The speeds of the two satellites will be in the ratio. 1999M

10. The escape velocities on two planets of masses	20. The acceleration due to gravity on the surface of
m_1 and m_2 and having same radius are v_1 and v_2	moon is 1/6 that on the earth and the diameter of
respectively then	the earth is 4 times the diameter of the moon. The ratio of the escape velocity of the moon to that of
1) $\frac{v_1}{v_2} = \frac{m_1}{m_2}$ 2) $\frac{v_2}{v_1} = \frac{m_1}{m_2}$	the earth is
$1) v_2 m_2 \qquad 2) v_1 m_2$	1) 1:4 2) 4:1 3) 5:1 4) 1:5
v_{1} $(m_{1})^{2}$ v_{1} (m_{1})	1992E 21. Two satellites A and B go round the earth in ciruclar
3) $\frac{v_1}{v_2} = \left(\frac{m_1}{m_2}\right)^2$ 4) $\frac{v_1}{v_2} = \sqrt{\frac{m_1}{m_2}}$	orbits at a height of R_A and R_B respectively from the
1998E	surface of the earth. Assume the earth to be a uniform
11. A satellite is revolving near the earth's surface. Its	sphere of radius R_{E} . The ratio of the magnitudes of the value of
orbital velocity is	the velocities of the satellites V_A/V_B is
1) 5.8 km/s 2) 18.4 km/s 2) 11.2 km/s	$1) \sqrt{\frac{R_B}{R_A}} 2) \frac{R_B + R_E}{(R_A + R_E)} 3) \sqrt{\frac{(R_B + R_E)}{(R_A + R_E)}} 4) \left(\frac{R_A}{R_B}\right)^2$
3) 11.2 km/s 4) 8.0 km/s 1998M	
12. The ratio of the escape velocity and the orbital	1991E
velocity is	22. If the escape velocity on earth is 11.2 km/sec, its value for a planet having double the radius and 8
1) $\sqrt{2}$ 2) $\frac{1}{\sqrt{2}}$ 3) 2 4) $1/2$	times the mass of earth is m./sec.
1) $\sqrt{2}$ 2) $\sqrt{2}$ 3) 2 4) 1/2	1990E
1998M	23. The escape velocity from the earth for a rocket is 11.2 km/sec. Ignoring the air resistance, the escape
13. The escape velocity of a sphere of mass 'm' is	velocity of 10 mg grain of sand from the earth will
given by $2GM$ $2GM$ $2GM$	be 1) 0.112 hm (see
1) $\sqrt{\frac{2GMm}{R_e}}$ 2) $\frac{2GM}{R_e^2}$ 3) $\sqrt{\frac{2GMm}{R_e^2}}$ 4) $\sqrt{\frac{2GM}{R_e}}$	1) 0.112 km/sec 2) 11.2 km/sec 3) 1.12 km/sec 4) None
$\begin{array}{c} \gamma K_e \gamma K_e \gamma K_e \gamma K_e \\ \end{array}$	1989E
14. The radius vector drawn from the sun to a planet	24. The orbital period of oscillation of an artificial
sweeps out areas in equal time	satellite revolving in a geostationary orbit is 1988E
1) equal 2) unequal 3) greater 4) less	25. If the earth is at one-fourth of its present distance
1996E 15. The escape velocity of a body from the earth is u.	from the sun, the duration of the year would be
What is the escape velocity from a planet whose	 Half the present year One-eigth the present year
mass and radius are twice those of the earth?	3) One-fourth the present year
1)2u 2)u 3)4u 4)16u 1995 E	4) One -sixteenth the present year
16. The orbital speed for an earth satelite near the	1984E 26. The mass of the earth is 9 times that of Mars. The
surface of the earth is 7 km/sec. If the radius of	radius of the earth is twice that of Mars. The
the orbit is 4 times the radius of the earth, the orbital	escape velocity of the earth is 12 km/sec. The
speed would be 1) 3.5 km/sec 2) 7 km/sec	escape velocity on Mars is km/sec. KEY
$3) 7 \sqrt{2}$ km/sec 4) 14 km/sec	1.4 2.4 3.2 4.4 5.2
1994E	6.2 7.1 8.4 9.2 10.4
17. A geostationary satellite has an orbital period of	$\begin{array}{ c c c c c c c c c c c c c c c c c c c$
1) 2 hours 2) 6 hours 3) 24 hours 4) 12 hours	$\begin{array}{cccccccccccccccccccccccccccccccccccc$
18. The escape velocity of an object on a planet whose radius is 4 times that of the earth and 'g' value 9	24.24hrs 25.2 26.4 $\sqrt{2}$
times that on the earth, in km.s ⁻¹ , is	
1) 33.6 2) 67.2 3) 16.8 4) 25.2	
10 The set all the investment of the set of	
19. The satellite is orbiting a planet at a certain height in a circular orbit. If the mass of the planet is	
reduced to half, the satellite would 1993M	
1) fall on the planet2) go to orbit of smaller radius	
3) go to orbit of higher radius	
4) escape from the planet	
JR.PHYSICS	270 GRAVITATION

QUESTIONS FROM OTHER COMPETETIVE EXAMS

1. A satellite of mass m revolves around the earth of radius R at a height x from its surface. If g is the acceleration due to gravity on the surface of the earth, the orbital speed of the satellite is

[AIEEE-2004]

1) gx 2)
$$\left(\frac{gR^2}{R+x}\right)^{1/2}$$
 3) $\frac{gR^2}{R+x}$ 4) $\frac{gR}{R-x}$

- 2. If 'g' is acceleration due to gravity on the earth's surface, the gain in the potential energy of an object of mass 'm' raised from the surface of the earth to a height equal to the radius 'R' of the earth is [AIEEE -2004]
- 1) 2 mgR 2) mgR 3) mgR/4 4) mgR/2
 Suppose the gravitational force varies inversely as the nth power of distance, then the time period of a planet in circular orbit of radius 'R' around the sun will be proportional to [AIEEE -2004]

1) $R^{\left(\frac{n+1}{2}\right)}$ 2) $R^{\left(\frac{n-2}{2}\right)}$ 3) R^{n} $R^{\left(\frac{n-1}{2}\right)}$ 4. The escape velocity for a body projected vertically upwards from the surface of earth is 11 km/s. If the body is projected at an angle of 45° with the vertical, the escape velocity will be [AIEEE -2003]

1) $11\sqrt{2}$ km/s 2) 22 km/s

3) 11 km/s 4) $11\sqrt{2} \text{ km/s}$

- 5. The time period of satellite of earth is 5 hr. If the separation between earth and the satellite is increased to 4 times the previous value, the new time period will become. [AIEEE -2003]
- 1) 10 hours 2) 80 hours 3) 40 hours 4) 20 hours
 A geo-stationary satellite orbits around the earth in a circular orbit of radius 36000 km. Then, the period of spy satellite orbiting a few hundred kilometers above the earth's surface (R_{earth} = 64000 km) will become.

[IIT Screening 2002]

1) (1/2) hr 2) 1 hr 3) 2 hr 4) 4 hr

7. The ratio of the radii of planets A and B is K_1 and ratio of accelerations due to gravity on them is K_2 The ratio of escape velocities from them will be [BHU -2002]

1)
$$K_1 K_2 = 2 \sqrt{K_1 K_2} 3 \sqrt{\frac{K_1}{K_2}} = 4 \sqrt{\frac{K_2}{K_1}}$$

8. The kinetic energy needed to project a body of mass m from earth's surface (radius R) to infinity is [AIEEE -2002]

$$1)\frac{mgR}{2} = 2)2mgR = 3) mgR = 4)\frac{mgR}{4}$$

9. The escape velocity of a body depends upon its mass as [AIEEE -2002]

1) m^0 2) m^1 3) m^3 4) m^2 10. Energy required to move a body of mass 'm' from an orbit of radius 2R to 3R is [AIEEE -2002]

1)
$$\frac{GMm}{2R^2}$$
 2) $\frac{GMm}{3R^2}$ 3) $\frac{GMm}{8R}$ 4) $\frac{GMm}{6R}$

11. If suddenly the gravitational force of attraction between earth and satellite revolving around it becomes zero, then the satellite will

[AIEEE -2002]

1) Continue to move in its orbit with same velocity

2) Move tangential to the original orbit with the same velocity

3) Becomes sationary in its orbit

4) Move towards the earth

12. The angular velocity of rotation of a star (of mass M and radius R) at which the matter will start escaping from its equator is

1)
$$\sqrt{\frac{2GR}{M}}$$
 2) $\sqrt{\frac{2GM}{R^3}}$ 3) $\sqrt{\frac{2GM}{R}}$ 4) $\sqrt{\frac{2GM^2}{R}}$

- 13. The time of revolution of planet A around the sun is 8 times that of another planet B. The distance of planet A from the sun is how many times greater than that of the planet B from the sun [AIEEE -2002]
 1) 2 2) 3 3)4 4) 5
- 14. The distance of Neptune and saturn from the Sun are respectively. 10^{13} and 10^{12} meters and their periodic times are respectively T_n and T_s . If their orbits are assumed to be circular, the value of T_n / T_s is [MP PMT-2001]

1)100 2)
$$10\sqrt{10}$$
 3) $\frac{1}{10\sqrt{10}}$ 4)10

15. Assume that the acceleration due to gravity on the surface of the moon is 0.2 times the acceleration due to gravity on the surface of the earth. If

 R_e is the maximum, range of a projectile on the earth's surface, what is the maximum range on the surface of the moon for the same velocity of projection [Kerala PET-2001]

1)
$$0.2R_e$$
 2) $2R_e$ 3) $0.5R_e$ 4) $5R_e$

16. The period of revolution of an earth's satellite close to the surface of earth is 90 minutes. The period of another earth's satellite in an orbit at a distance of three times earth's radius from its surface will be [J&K CET-2001]

1) 90 minutes'2) $90 \times \sqrt{8}$ minutes3) 270 minutes4) 720 minutes

17If the spinning speed of earth is increased then
the weight of the body at equator
[UP CP MTE2000,AFMC-2002]
1) Does notchange 2) Doubles
3) Decreases 4) Increases
(Aspace craft is launched in a circular orbit trong the
earth's gravitational pull circMT 2001]
1) 20.2
$$K_{MS}^{-1}$$

2) 3, $25KmS^{-1}$
2) 3, $8KmS^{-1}$
2) 3, $8KmS^{-1}$
2) 3, $8KmS^{-1}$
2) $3KmS^{-1}$
2) Reparent motion of Sun round the earth
directed towards Sun?
2) Phenomenon of day and night
2) Phenomenon of day and night
2) Phenomenon of the falling body towards earth
4) Deviation of the falling body towards earth
4) Deviation of the falling body towards earth
orbit of readus R is T. Its period of revolution in
an orbit of radius R is T. Its period of revolution in
at a distance X from the carter is: (CHNI 1997)
1) $33 KmS^{-2}$
2) $11/\sqrt{3} KmS^{-1}$
2). $11/\sqrt{3} KmS^{-1}$
2) The period of revolution of acter the sum
at a distance X from the carter is: (CHNI 1997)
1) $33 KmS^{-2}$
2) $11/\sqrt{3} KmS^{-1}$
2) $11/\sqrt{3} KmS^{-1}$
2) $11/\sqrt{3} KmS^{-1}$
3) $11/\sqrt{3} KmS^{-1}$
3) $11/\sqrt{3} KmS^{-1}$
3) $11/\sqrt{3} KmS^{-1}$
2) $11/\sqrt{3} KmS^{-1}$
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3) $11/\sqrt{3} KmS^{-1}$
3